# Ab Initio Calculations of Amplitude and Phase Functions for Extended X-ray Absorption Fine Structure Spectroscopy 

Boon-Keng Teo* and P. A. Lee<br>Contribution from Bell Laboratories, Murray Hill, New Jersey 07974.<br>Received October 16, 1978


#### Abstract

Extended X-ray absorption fine structure (EXAFS) amplitude and phase functions have been calculated from first principle for $K$ and $L$ edges of nearly half of the elements in the periodic table. It is shown that for $L_{11,111}$ edges, the transition from the initial state to the $d$ final state is favored by almost a factor of 50 over the transition to the s final state such that the $L_{11.111}$ EXAFS can be analyzed in the same way as $K$ and $L_{1}$ edges with the use of the $l=2$ phase shifts. These theoretical EXAFS functions exhibit significant new structures (as a function of electron wave vector) which are in accord with experiments. Chemically interesting trends are observed for these functions as a function of atomic number $Z$. It is believed that these ab initio EXAFS functions can be used in EXAFS data analysis to provide accurate structural (interatomic distances) and chemical (type and number of neighboring atoms, Debye-Waller factors) information.


## Introduction

The phenomenon of EXAFS refers to the oscillatory modulation of the X-ray absorption coefficient as a function of X-ray photon energy beyond the absorption edge. The existence of such an extended fine structure has been known and treated theoretically in the 1930s by Kronig. ${ }^{1}$ Recent developments ${ }^{2-21}$ initiated by the work of Sayers, Stern, and Lytle ${ }^{2-5}$ have led to the recognition of the structural content of this technique. At the same time, the availability of synchrotron radiation has greatly improved the speed of data acquisition and $S / N$ by a factor of $\sim 10^{5}$ over conventional X-ray sources. ${ }^{11-13}$ Such developments have gradually established EXAFS as a practical structural tool.

It is now generally accepted that the EXAFS phenomenon is due to a final state interference effect involving scattering of the outgoing photoelectron from the neighboring atoms. This causes an oscillatory behavior of the absorption rate. For reasonable high energy and moderate thermal vibrations, the modulation of the absorption coefficient, normalized to the "background" absorption ( $\mu_{0}$ ), can be described in terms of the scattering amplitude from the neighboring atom $F(k)$ and the phase shift function $\phi(k)$ which consists of contribution from the absorbing atom and the neighboring atom. If these functions are known, it should be possible to deduce structural information about the local environment of the absorber.

Two major approaches of data analysis have since been developed: the Fourier transform ${ }^{2-8}$ and the curve-fitting techniques. ${ }^{14-21}$ Both of these methods require a detailed knowledge of the amplitude and phase functions for the determination of chemical information such as coordination number, Debye-Waller factor, and interatomic distances through the assumptions of amplitude and phase transferabilities. These latter hypotheses, which greatly enhance the chemical content of EXAFS spectroscopy, have been demonstrated previously. ${ }^{4}$.10.17.21

Experimentally the amplitude and phase functions can be obtained from the EXAFS spectra of model compounds ${ }^{2-5.9-11.14,19-21}$ However, it is often possible to obtain only the product (vide infra) $F_{b}(k) e^{-2 \pi^{2} k^{2}} e^{-2 r / \lambda}$ (assuming $N$ is known) for each type of scatterer B and the combined phase $\phi_{a b}(k)$ for each pair of atoms AB (assuming $r$ is known) from experimental data. The extraction of amplitude function $F_{b}(k)$ alone requires knowledge of the Debye-Waller factor $\sigma$ (from a separate study or from temperature-dependent measurements) and the electron mean free path $\lambda$ whereas the separation of the total phase shift $\phi_{a b}$ into individual phases $\phi_{a}$ (due to the absorber A) and $\phi_{b}$ (due to the backscatterer B) can only be achieved by measuring the phase shifts of var-
ious combinations of pairs of atoms and arbitrarily defining $\phi_{a}$ or $\phi_{b}$ for one atom. ${ }^{22}$

To avoid the tedious task of searching, measuring, and analyzing model compounds, it is clearly desirable to calculate the amplitude $F(k)$ and the individual phase shifts $\phi_{a}(k)$ and $\phi_{b}(k)$ from first principle. With an accurate method, this not only represents a major saving in time and effort, but also greatly reduces the danger of introducing experimental errors from the analyses of model compounds. ${ }^{15-20}$

We have calculated the amplitude and phase functions of nearly half of the elements in the periodic table using an electron-atom scattering theory recently introduced by Lee and Beni. ${ }^{8}$ In this paper, we tabulate and plot our results as a function of atomic number $Z$ such that the intermediate elements can readily be interpolated. We have also tested some of these functions in EXAFS data analysis with favorable results. ${ }^{15-20}$ Chemically interesting trends can be observed from these plots which either help clarify questions in EXAFS spectroscopy or caution ways of interpreting EXAFS data. The use of these theoretical functions in conjunction with Fourier transform, curve fitting, or a combination of both in EXAFS data analyses has been proven to be highly valuable, especially for complex multiatom, multidistance systems. ${ }^{16.18-20}$

## EXAFS of $K$ and $L$ Edges

For the excitation of a s level ( $K$ or $L_{1}$ edge) the absorption coefficient normalized to a smooth background $\mu_{0}$ can be described by

$$
\begin{align*}
& \chi(k)=\frac{\mu-\mu_{0}}{\mu_{0}} \\
& \quad=\sum_{j} 3 \cos ^{2} \theta_{j} N_{j} F_{j}(k) e^{-2 \pi \sigma_{j}^{2} k^{2}} e^{-2 r_{j} / \lambda} \frac{\sin \left(2 k r_{j}+\phi_{j}(k)\right)}{k r_{j}^{2}} \tag{1}
\end{align*}
$$

where $F_{j}(k)$ is the backscattering amplitude from each of the $N_{j}$ neighboring atoms of the $j$ th type with a Debye-Waller factor of $\sigma_{j}$ (to account for thermal vibration and static dis-order-viz., nonequivalent distances) and at a distance $r_{j}$ away. The $j$ th neighbor makes an angle $\theta_{j}$ with the polarization vector of the X -ray and the factor $3 \cos ^{2} \theta_{j}$ averages to $I$ for polycrystalline samples. $\phi_{j}(k)$ is the total phase shift experienced by the photoelectron and is given by

$$
\begin{equation*}
\phi(k)=\phi_{a}^{l=1}(k)+\phi_{b}(k)-\pi \tag{2}
\end{equation*}
$$

where $\phi_{b}$ is the phase of the backscattering amplitude from the neighbor and $\phi_{a}{ }^{\prime}$ equals $2 \delta_{l}^{\prime}$ where $\delta_{l}{ }^{\prime}$ is the $l$ phase shift due to the central atom. ${ }^{8}$ The term $e^{-2 r_{j} / \lambda}$ is due to inelastic


Figure 1. The ratio of the radial dipole matrix elements $M_{21} / M_{01}$ (where $M_{21}$ and $M_{01}$ correspond to the $l=1$ initial atomic state and the $l=2$ and $l=0$ final states, respectively) as a function of electron wave vector $k$ for $\mathrm{Ti}, \mathrm{Zr}$, and W .
(scattering) losses with $\lambda$ being the electron mean free path. Finally $k$ refers to the photoelectron wave vector which is defined as

$$
\begin{equation*}
k=\sqrt{\frac{2 m}{\hbar^{2}}\left(E-E_{0}\right)} \tag{3}
\end{equation*}
$$

where $E$ is the photon energy and $E_{0}$ is the energy threshold of the absorption edge. It should also be noted that, while the backscattering amplitude is a function of the scatterer only, the phase shift depends upon both the absorber (central atom) and the backscatterer (neighboring atom).

The description of EXAFS from exciting a p level ( $L_{11,111}$ spectra) is complicated by the fact that the initial $p$ state can go to a final state of $s$ or d symmetry. Instead of the single term in eq 1 we have three terms (eq 4) in which the central atom phase $\phi_{a}$ is given by $2 \delta_{2}{ }^{\prime}, 2 \delta_{0}{ }^{\prime}$, and $\delta_{0}{ }^{\prime}+\delta_{2}{ }^{\prime}$ where $\delta_{1}^{\prime}$ is the phase shift of an outgoing wave with angular momentum l. ${ }^{7.23}$

$$
\begin{align*}
& \chi(k)= \sum_{j} \frac{N_{j} F_{j}(k)}{k r_{j}^{2}} e^{-2 \sigma_{j} k^{2} e^{-2 r_{j} / \lambda}} \\
& \times\left\{1 / 2\left(1+3 \cos ^{2} \theta_{j}\right)\left|M_{21}\right|^{2} \sin \left(2 k r_{j}+\phi_{2 j}\right)\right. \\
&+1 / 2\left|M_{01}\right|^{2} \sin \left(2 k r_{j}+\phi_{0 j}\right) \\
&\left.+M_{01} M_{21}\left(1-3 \cos ^{2} \theta_{j}\right) \sin \left(2 k r_{j}+\phi_{02 j}\right)\right\} \\
& \times\left(\left|M_{21}\right|^{2}+1 / 2\left|M_{01}\right|^{2}\right)^{-1} \tag{4}
\end{align*}
$$

The phase functions for the $j$ th atom are

$$
\begin{gather*}
\phi_{2}=\phi_{a}^{l=2}+\phi_{b}  \tag{5a}\\
\phi_{0}=\phi_{a}^{l=0}+\phi_{b}  \tag{5b}\\
\phi_{02}=1 / 2\left(\phi_{a}^{l=2}+\phi_{a}^{l=0}\right)+\phi_{b} \tag{5c}
\end{gather*}
$$

Note that in comparison with the $K$ edge or $L_{1}$ edge phase given by eq 2 the additional factor of $\pi$ is absent in eq 5 . This is because the $\pi$ factor has been introduced to take care of an overall minus sign which is absent for $L_{11,111}$ edge (viz., $\chi(k)$ has a $(-1)^{\prime}$ factor where $l$ is the angular momentum of the initial state). The matrix elements $M_{01}$ and $M_{21}$ are the radial dipole matrix elements between the $2 \mathrm{p}(l=1)$ atomic wave function and the $l=2$ and $l=0$ final states.

For polycrystalline samples the cross term involving $\delta_{0}{ }^{\prime}+$ $\delta_{2}^{\prime}$ (i.e., the third term in braces in eq 4) vanishes by angular averaging. We are still left with a complicated expression which requires two sets of central atom phase shifts and the ratio between the dipole matrix elements. In our calculation of the phase shift (to be discussed later) we calculate the final state wave function of angular momentum $l$ as an intermediate step.

It is straightforward to compute the matrix elements $M_{21}$ and $M_{01}{ }^{24}$ The ratio $M_{21} / M_{01}$ is plotted in Figure 1 as a function of photoelectron momentum $k$ for $\mathrm{Ti}, \mathrm{Zr}$, and W . We see that the ratio is of the order of 5 and is relatively independent of $k$. For light atoms, however, $M_{21} / M_{01}$ shows some $k$ dependence and can vary by as much as a factor of 2 as in, for example, chlorine. From eq 4 we see that transitions to the $d$ final state are favored by a factor of 50 . Thus for all practical purposes $M_{01}$ can be ignored and the $L_{11,11 t}$ edge can be analyzed in the same way as $K$ and $L_{1}$ edges with the use of the $l=2$ phase shift given by eq 5 a in place of eq 2 .

It is interesting to note that the calculated ratio is in excellent agreement with the value $M_{01} / M_{21}=0.2 \pm 0.06$ obtained by Heald and Stern ${ }^{23}$ from the angular dependence of the tungsten $L_{111}$ edge spectrum. The physical reason that $M_{01}$ is smaller is that the $l=0$ final state must be orthogonal to the Is core state and therefore is much more rapidly oscillatory than the $l=2$ final state in the region of the $2 p$ wave function. It should also be mentioned that Lytle, Sayers, and Stern ${ }^{25}$ have observed sidebands in the Fourier transform spectrum in the gold $L$ edge and interpreted them as arising from the $l$ $=0$ final states. The smallness of ratio $M_{01} / M_{21}$ means that the $l=0$ contribution is practically unobservable and that the sideband is more properly interpreted as arising from the complicated structure in the amplitude and phase functions for heavy atoms (vide infra).

## Methods

The ab initio EXAFS amplitude and phase functions presented in this work were calculated with the electron-atom scattering theory originally developed by Lee and Beni. ${ }^{8} \mathrm{Ba}$ sically the theory involves the construction of an effective complex scattering potential that adequately accounts for the exchange and correlation effects caused by the electrons in the atom using a modified Thomas-Fermi approach which amounts to replacing the atom by an electron gas with spatially varying density and calculating the self-energy using the plasmon pole approximation. Specifically, the method amounts to first calculating the spatial dependence of the charge density from Hartree-Fock wave functions (vide infra). From this the local Fermi energy and the local momentum can be obtained by using the Thomas-Fermi description of an atom. Then a complex potential is constructed which depends on the kinetic energy of the incoming electron. This complex potential is added to the electrostatic potential to yield the complex phase shifts. This scheme, when applied to the electron-atom scattering problem, gives the backscattering amplitude $F_{b}(k)$ and the backscattering phase shift $\phi_{h}(k)$ for the scatterer and the central atom phase shift $\phi_{a}{ }^{\prime}(k)$ for the absorber in EXAFS spectroscopy. The results for nearly half of the elements in the periodic table with atomic number $Z<86$ are tabulated in Tables 1-V111.

Tables 1-111 were calculated using Clementi-Roetti wave functions. ${ }^{26 a}$ The central atom phase shifts $\phi_{a}{ }^{\prime}(k)$ were obtained using the unscreened $Z+1$ ion approximation ${ }^{8}$ which amounts to using the $Z+1$ atomic wave function with one outer (valence) electron missing. The latter was chosen as a compromise for a screened $Z+1$ atom (completely relaxed case) at low kinetic energy and an unscreened $Z$ ion (completely unrelaxed) at high kinetic energy. These phase shifts have previously been parametrized and published. ${ }^{17}$ We have extended these calculations to heavier atoms. We find it more convenient to use the Herman-Skillman wave functions. ${ }^{26 \mathrm{~b}}$ Furthermore, for heavier atoms the backscattering amplitude and phase functions cannot be parametrized by simple analytical functions. These more recent results are tabulated in Tables IV-VIII. For the central atom phase shifts the $Z$ ion with one core electron (1s or 2s) missing was used. In all backscattering amplitude (Tables I and IV) and phase (Tables

Table I. Backscattering Amplitudes $F(k)$ in $\AA$ vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Clementi-Roetti Wave Functions

| Z | CHEM | $\begin{array}{r} \mathrm{k}=3.7795 \\ 8.5038 \end{array}$ | $\begin{aligned} & 4.2519 \\ & 9.4486 \end{aligned}$ | $\begin{array}{r} 4.7243 \\ 10.3935 \end{array}$ | $\begin{array}{r} 5.1967 \\ 11.3384 \end{array}$ | $\begin{array}{r} 5.6692 \\ 12.2832 \end{array}$ | $\begin{array}{r} 6.1416 \\ 13.2281 \end{array}$ | $\begin{array}{r} 6.6140 \\ 14.1729 \end{array}$ | $\begin{array}{r} 7.0865 \\ 15.1178 \end{array}$ | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 6 | C | 0.5063 | 0.3566 | 0.2800 | 0.2430 | 0.2099 | 0.1719 | 0.1430 | 0.1325 | 0.1182 |
|  |  | 0.0839 | 0.0712 | 0.0562 | 0.0481 | 0.0397 | 0.0349 | 0.0297 | 0.0267 |  |
| 8 | 0 | 0.7041 | 0.5781 | 0.4629 | 0.3909 | 0.3248 | 0.2661 | 0.2205 | 0.1913 | 0.1700 |
|  |  | 0.1256 | 0.1016 | 0.0814 | 0.0674 | 0.0563 | 0.0489 | 0.0416 | 0.0365 |  |
|  | F | 0.7175 | 0.6275 | 0.5294 | 0.4503 | 0.3924 | 0.3273 | 0.2733 | 0.2349 | 0.2038 |
|  |  | 0.1532 | 0.1204 | 0.0956 | 0.0792 | 0.0652 | 0.0566 | 0.0481 | 0.0423 |  |
| 11 | NA | 0.6630 | 0.6007 | 0.5436 | 0.4669 | 0.4310 | 0.3905 | 0.3482 | 0.3023 | 0.2657 |
|  |  | 0.2028 | 0.1610 | 0.1263 | 0.1014 | 0.0864 | 0.0715 | 0.0629 | 0.0569 |  |
| 15 | P | 0.7829 | 0.7034 | 0.6517 | 0.6112 | 0.5533 | 0.4806 | 0.4471 | 0.4164 | 0.3671 |
|  |  | 0.2938 | 0.2344 | 0.1874 | 0.1571 | 0.1295 | 0.1089 | 0.0939 | 0.0845 |  |
| 16 | S | 0.8140 | 0.8173 | 0.7569 | 0.6523 | 0.5843 | 0.5505 | 0.5075 | 0.4508 | 0.3936 |
|  |  | 0.3283 | 0.2608 | 0.2151 | 0.1731 | 0.1459 | 0.1191 | 0.1016 | 0.0854 |  |
| 17 | CL | 0.8269 | 0.8413 | 0.8188 | 0.7130 | 0.6344 | 0.5755 | 0.5391 | 0.4938 | 0.4361 |
|  |  | 0.3475 | 0.2864 | 0.2289 | 0.1901 | 0.1580 | 0.1330 | 0.1100 | 0.0954 |  |
| 20 | CA | 0.6478 | 0.8064 | 0.8361 | 0.7698 | 0.7051 | 0.6457 | 0.5971 | 0.5404 | 0.4972 |
|  |  | 0.3981 | 0.3317 | 0.2776 | 0.2160 | 0.1910 | 0.1630 | 0.1359 | 0.1146 |  |
| 22 | TI | 0.6225 | 0.7179 | 0.7580 | 0.8075 | 0.8073 | 0.7404 | 0.6822 | 0.6198 | 0.5537 |
|  |  | 0.4707 | 0.3844 | 0.3378 | 0.2653 | 0.2197 | 0.1905 | 0.1587 | 0.1391 |  |
| 24 | CR | 0.4569 | 0.5482 | 0.6666 | 0.7161 | 0.7604 | 0.7668 | 0.7479 | 0.6865 | 0.6354 |
|  |  | 0.5288 | 0.4438 | 0.3661 | 0.3089 | 0.2541 | 0.2195 | 0.1863 | 0.1623 |  |
| 26 | FE | 0.3625 | 0.4285 | 0.5148 | 0.5849 | 0.6800 | 0.7274 | 0.7201 | 0.6942 | 0.6615 |
|  |  | 0.5798 | 0.4949 | 0.4189 | 0.3474 | 0.2902 | 0.2451 | 0.2086 | 0.1768 |  |
| 29 | CU | 0.2757 | 0.2577 | 0.3559 | 0.4352 | 0.5124 | 0.6240 | 0.6680 | 0.6723 | 0.6682 |
|  |  | 0.6325 | 0.5587 | 0.4846 | 0.4113 | 0.3480 | 0.2951 | 0.2513 | 0.2178 |  |
| 32 | GE | 0.3062 | 0.2491 | 0.2677 | 0.3006 | 0.3833 | 0.4995 | 0.5529 | 0.5724 | 0.5982 |
|  |  | 0.6124 | 0.5788 | 0.5249 | 0.4674 | 0.3988 | 0.3458 | 0.2934 | 0.2543 |  |
| 35 | BR | 0.3836 | 0.3144 | 0.2785 | 0.2894 | 0.3386 | 0.4210 | 0.5020 | 0.5485 | 0.5694 |
|  |  | 0.6120 | 0.5832 | 0.5513 | 0.4975 | 0.4456 | 0.3826 | 0.3362 | 0.2899 |  |

Table II. Backscattering Phase Shifts $\phi_{b}(k)$ in Radian vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Clementi-Roetti Wave Functions

| Z | CHEM | $\begin{array}{r} \mathrm{k}=3.7795 \\ 8.5038 \end{array}$ | $\begin{aligned} & 4.2519 \\ & 9.4486 \end{aligned}$ | $\begin{array}{r} 4.7243 \\ 10.3935 \end{array}$ | $\begin{array}{r} 5.1967 \\ 11.3384 \end{array}$ | $\begin{array}{r} 5.6692 \\ 12.2832 \end{array}$ | $\begin{array}{r} 6.1416 \\ 13.2281 \end{array}$ | $\begin{array}{r} 6.6140 \\ 14.1729 \end{array}$ | $\begin{array}{r} 7.0865 \\ 15.1178 \end{array}$ | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 6 | C | 0.0481 | -0.0992 | -0.3146 | -0.5489 | -0.7327 | -0.9119 | -1.0949 | -1.2448 | -1.3145 |
|  |  | -1.6305 | -1.8127 | -2.0773 | -2.2421 | -2.4449 | -2.5762 | -2.7436 | -2.8657 |  |
| 8 | 0 | 0.5056 | 0.4555 | 0.3436 | 0.1858 | 0.0530 | $-0.0643$ | -0.1986 | -0.3627 | -0.5035 |
|  |  | -0.7327 | -0.9763 | -1.1969 | -1.4081 | -1.6035 | -1.7800 | -1.9252 | -2.0925 |  |
| 9 | F | 0.6436 | 0.5801 | 0.5072 | 0.3723 | 0.2630 | 0.1815 | 0.0750 | -0.0616 | -0.1872 |
|  |  | -0.4056 | -0.6354 | -0.8457 | -1.0587 | -1.2427 | -1.4357 | -1.5878 | -1.7465 |  |
| 11 | NA | 1.4871 | 1.4189 | 1.2842 | 1.1232 | 1.0379 | 0.9245 | 0.8037 | 0.7278 | 0.6167 |
|  |  | 0.4132 | 0.1683 | -0.0389 | -0,2383 | -0.3511 | -0.4625 | -0.8664 | -1.0499 |  |
| 15 | P | 3.5897 | 3.4657 | 3.3259 | 3.1419 | 3.0078 | 2.8527 | 2.7088 | 2.5551 | 2.4108 |
|  |  | 2.1811 | 1.9250 | 1.7299 | 1.4634 | 1.2613 | 1.0318 | 0.8387 | 0.6943 |  |
| 16 | S | 3.6618 | 3.5265 | 3.4586 | 3.3502 | 3.1962 | 3.0299 | 2.8992 | 2.7921 | 2.6714 |
|  |  | 2.4337 | 2.1703 | 1.9651 | 1.7471 | 1.5259 | 1.3306 | 1.1155 | 0.9559 |  |
| 17 | CL | 3.7781 | 3.6982 | 3.6428 | 3.5596 | 3.4464 | 3.3174 | 3.1796 | 3.0631 | 2.9656 |
|  |  | 2.7275 | 2.4810 | 2.2481 | 2.0427 | 1.8078 | 1.6304 | 1.4153 | 1.2463 |  |
| 20 | CA | 5.5111 | 5.1885 | 5.1244 | 5.0356 | 4.9437 | 4.7719 | 4.6259 | 4.4780 | 4.3182 |
|  |  | 4.0505 | 3.7842 | 3.5296 | 3.1526 | 2.9924 | 2.8367 | 2.6625 | 2.4975 |  |
| 22 | T1 | 5.0578 | 5.1829 | 5.0969 | 5.0310 | 5.0209 | 4.9533 | 4.8189 | 4.7412 | 4.6626 |
|  |  | 4.3738 | 4.1297 | 3.8988 | 3.6356 | 3.4189 | 3.1788 | 2.9330 | 2.6651 |  |
| 24 | CR | 4.6789 | 5.0195 | 5.1303 | 5.0781 | 5.0492 | 5.0313 | 5.0002 | 4.9126 | 4.8156 |
|  |  | 4.6585 | 4.4184 | 4.2252 | 3.9968 | 3.7829 | 3.5680 | 3.3656 | 3.1832 |  |
| 26 | FE | 4.6376 | 5.1013 | 5.2163 | 5.1940 | 5.1656 | 5.2196 | 5.1716 | 5.1143 | 5.0787 |
|  |  | 4.9115 | 4.7426 | 4.5374 | 4.3659 | 4.1507 | 3.9500 | 3.7421 | 3.5227 |  |
| 29 | CU | 4.3259 | 4.9690 | 5.3797 | 5.4192 | 5.4286 | 5.4584 | 5.4798 | 5.4447 | 5.3991 |
|  |  | 5.3159 | 5.1710 | 5.0326 | 4.8523 | 4.6898 | 4.4958 | 4.3136 | 4.1174 |  |
| 32 | GE | 4.4485 | 4.9937 | 5.4287 | 5.6053 | 5.6784 | 5.7439 | 5.7959 | 5.7799 | 5.7404 |
|  |  | 5.6405 | 5.5145 | 5.3850 | 5.2535 | 5.1026 | 4.9454 | 4.7666 | 4.6126 |  |
| 35 | BR | 4.6324 | 5.0076 | 5.5637 | 5.9716 | 6.2120 | 6.3555 | 6.3982 | 6.3956 | 6.3713 |
|  |  | 6.2729 | 6.1482 | 5.9926 | 5.8585 | 5.7174 | 5.5499 | 5.3929 | 5.2145 |  |

11 and $V$ ) function calculations, the wave functions were truncated at 1.5 times the covalent radius and a uniform charge density was added to preserve charge neutrality within this radius. For all central atom phase shift (Tables III, VI-VIII)
calculations, the Coulomb field was cut off at twice the covalent radius. All calculations were performed on neutral atoms except for a few alkali and alkali-earth elements, which were treated as cations. The difference between various oxidation

Table III. Central Atom Phase Shifts $\phi_{a}{ }^{l=1}(k)$ in Radian vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Clementi-Roetti Wave Functions and the $(Z+1)$ Ion Approximation

| Z | CHEM | $\begin{array}{r} \mathrm{k}=3.7795 \\ 8.5038 \end{array}$ | $\begin{aligned} & 4.2519 \\ & 9.4486 \end{aligned}$ | $\begin{array}{r} 4.7243 \\ 10.3935 \end{array}$ | $\begin{array}{r} 5.1967 \\ 11.3384 \end{array}$ | $\begin{array}{r} 5.6692 \\ 12.2832 \end{array}$ | $\begin{array}{r} 6.1416 \\ 13.2281 \end{array}$ | $\begin{array}{r} 6.6140 \\ 14.1729 \end{array}$ | $\begin{array}{r} 7.0865 \\ 15.1178 \end{array}$ | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 12 | MG | -3.5617 | -3.9685 | -4.3733 | -4.7701 | -5.0669 | -5.3543 | -5.6428 | -5.8338 | -6.0636 |
|  |  | -6.4312 | -6.7567 | -7.0415 | -7.2993 | -7.5158 | -7.7085 | -7.8942 | -8.0626 |  |
| 14 | S1 | -2.3921 | -2.9872 | -3.4091 | -3.7521 | -4.0891 | -4.4444 | -4.7279 | -4.9464 | -5.1959 |
|  |  | -5.6061 | -5.9793 | -6.2763 | -6.5604 | -6.8066 | -7.0308 | -7.2375 | -7.4235 |  |
| 16 | S | -1.5587 | -2.1042 | -2.5678 | -2.9541 | -3.2763 | -3.6236 | -3.9540 | -4.2141 | -4.4305 |
|  |  | -4.9052 | -5.2781 | -5.6170 | -5.9076 | -6.1779 | -6.4301 | -6.6363 | -6.8490 |  |
| 21 | SC | 1.1795 | 0.6230 | 0.1224 | -0.3994 | -0.7927 | -1.1724 | -1.5605 | -1.8614 | -2.1806 |
|  |  | -2.7211 | -3.2072 | -3.6470 | -4.0255 | -4.3628 | -4.6688 | -4.9459 | -5.2051 |  |
| 23 | v | 1.5350 | 0.9061 | 0.3468 | -0.0757 | -0.4409 | -0.8628 | -1.1871 | -1.4872 | -1.8112 |
|  |  | -2.3372 | -2.8357 | -3.2666 | -3.6301 | -3.9812 | -4.2905 | -4.5668 | -4.8298 |  |
| 26 | FE | 2.0856 | 1.4838 | 0.9136 | 0.4869 | 0.1332 | -0.2651 | -0.6151 | -0.9053 | -1.2124 |
|  |  | -1.7320 | -2.2309 | -2.6475 | -3.0347 | -3.3885 | -3.7033 | -3.9952 | -4.2651 |  |
| 29 | CU | 2.5785 | 1.9474 | 1.4064 | 1.0160 | 0.6236 | 0.2288 | -0.0929 | -0.3817 | -0.6966 |
|  |  | -1.2194 | -1.7082 | -2.1233 | -2.5115 | -2.8684 | -3.1826 | -3.4880 | -3.7629 |  |
| 32 | GE | 3.2853 | 2.6446 | 2.1916 | 1.7551 | 1.3016 | 0.9493 | 0.6306 | 0.2722 | -0.0379 |
|  |  | -0.5737 | -1.0568 | -1.5087 | -1.8904 | -2.2483 | -2.5956 | -2.8984 | -3.1768 |  |
| 35 | BR | 4.2011 | 3.5744 | 3.0770 | 2.6538 | 2.2224 | 1.8180 | 1.4768 | 1.1480 | 0.8028 |
|  |  | 0.2540 | -0.2749 | -0.7220 | -1.1320 | -1.5120 | -1.8640 | -2.1812 | -2.4812 |  |
| 40 | ZR | 7.1080 | 6.4124 | 5.8533 | 5.2006 | 4.7033 | 4.2814 | 3.8093 | 3.4067 | 3.0375 |
|  |  | 2.3621 | 1.7442 | 1.1970 | 0.7108 | 0.2704 | -0.1353 | -0,5099 | -0.8646 |  |
| 45 | RH | 6.4865 | 6.0246 | 5.5155 | 4.9999 | 4.5687 | 4.2071 | 3.8180 | 3.4570 | 3.1318 |
|  |  | 2.5292 | 1.9858 | 1.4785 | 1.0327 | 0.6200 | 0.2359 | -0.1138 | -0.4461 |  |
| 52 | TE | 9.1879 | 8.5796 | 8.0058 | 7.4207 | 6.9617 | 6.5146 | 6.0481 | 5.6693 | 5.2845 |
|  |  | 4.6178 | 3.9937 | 3.4310 | 2.9335 | 2.4710 | 2.0384 | 1.6423 | 1.2821 |  |

Table IV. Backscattering Amplitudes $F(k)$ in $\AA$ vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Herman-Skillman Wave Functions

| Z | CHEM | $\mathrm{k}=3.7795$ | 4.2519 | 4.7243 | 5.1967 | 5.6692 | 6.1416 | 6.6140 | 7.0865 | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 8.5038 | 9.4486 | 10.3935 | 11.3384 | 12.2832 | 13.2281 | 14.1729 | 15.1178 |  |
| 14 | S1 | 0.7326 | 0.7068 | 0.6358 | 0.5533 | 0.5208 | 0.4815 | 0.4223 | 0.3760 | 0.3439 |
|  |  | 0.2745 | 0.2215 | 0.1771 | 0.1466 | 0.1183 | 0.1004 | 0.0836 | 0.0727 |  |
| 17 | CL | 0.7740 | 0.8130 | 0.7927 | 0.7050 | 0.6256 | 0.5710 | 0.5316 | 0.4883 | 0.4323 |
|  |  | 0.3441 | 0.2857 | 0.2288 | 0.1905 | 0.1597 | 0.1338 | 0.1111 | 0.0960 |  |
| 20 | CA | 0.6805 | 0.7672 | 0.8078 | 0.7555 | 0.7002 | 0.6389 | 0.5916 | 0.5387 | 0.4937 |
|  |  | 0.3976 | 0.3310 | 0.2771 | 0.2213 | 0.1915 | 0.1618 | 0.1350 | 0.1135 |  |
| 40 | ZR | 0.6463 | 0.5510 | 0.4291 | 0.3454 | 0.3190 | 0.3513 | 0.4217 | 0.4807 | 0.5192 |
|  |  | 0.5842 | 0.5784 | 0.5596 | 0.5253 | 0.4791 | 0.4349 | 0.3904 | 0.3366 |  |
| 42 | MO | 0.7500 | 0.6977 | 0.5944 | 0.4233 | 0.3582 | 0.3596 | 0.3894 | 0.4570 | 0.5172 |
|  |  | 0.5779 | 0.5960 | 0.5729 | 0.5452 | 0.5019 | 0.4509 | 0.4085 | 0.3668 |  |
| 44 | RU | 0.8645 | 0.7911 | 0.7219 | 0.5553 | 0.4243 | 0.3835 | 0.3894 | 0.4305 | 0.4903 |
|  |  | 0.5732 | 0.5976 | 0.5900 | 0.5623 | 0.5218 | 0.4755 | 0.4239 | 0.3837 |  |
| 46 | PD | 0.9115 | 0.8554 | 0.7922 | 0.6655 | 0.5083 | 0.4197 | 0.3822 | 0.3984 | 0.4530 |
|  |  | 0.5519 | 0.5944 | 0.5994 | 0.5813 | 0.5413 | 0.4974 | 0.4470 | 0.4068 |  |
| 47 | AG | 0.9143 | 0.8729 | 0.8121 | 0.6993 | 0.5558 | 0.4354 | 0.3745 | 0.3895 | 0.4358 |
|  |  | 0.5330 | 0.5923 | 0.6008 | 0.5881 | 0.5517 | 0.5066 | 0.4616 | 0.4198 |  |
| 50 | SN | 1.1146 | 1.0135 | 0.8812 | 0.7798 | 0.6296 | 0.4939 | 0.4107 | 0.4014 | 0.4261 |
|  |  | 0.5174 | 0.5849 | 0.6142 | 0.6031 | 0.5789 | 0.5380 | 0.4940 | 0.4395 |  |
| 53 | 1 | 1.1468 | 1.0382 | 0.9153 | 0.8085 | 0.6808 | 0.5630 | 0.4660 | 0.4215 | 0.4238 |
|  |  | 0.5065 | 0.5883 | 0.6233 | 0.6313 | 0.6063 | 0.5680 | 0.5190 | 0.4749 |  |
| 57 | LA | 0.6221 | 0.7004 | 0.7250 | 0.7630 | 0.6704 | 0.5967 | 0.5238 | 0.4512 | 0.4503 |
|  |  | 0.5067 | 0.5901 | 0.6420 | 0.6381 | 0.6242 | 0.6089 | 0.5548 | 0.5164 |  |
| 58 | CE | 0.6161 | 0.6850 | 0.6812 | 0.7354 | 0.6726 | 0.5940 | 0.5272 | 0.4443 | 0.4224 |
|  |  | 0.4573 | 0.5524 | 0.6158 | 0.6300 | 0.6188 | 0.6083 | 0.5663 | 0.5188 |  |
| 65 | TB | 0.3467 | 0.4390 | 0.4390 | 0.5228 | 0.5789 | 0.5519 | 0.5369 | $0.4821$ | 0.4263 |
|  |  | 0.3606 | 0.3836 | 0.4589 | 0.5367 | 0.5707 | 0.6010 | 0.5999 | 0.5772 |  |
| 70 | YB | 0.2904 | 0.3517 | 0.3241 | 0.3897 | 0.4710 | 0.4782 | 0.4918 | 0.4765 | 0.4360 |
|  |  | 0.3677 | 0.3169 | 0.3367 | 0.4202 | 0.4868 | 0.5316 | 0.5826 | 0.5791 |  |
| 74 | W | 0.3196 | 0.2523 | 0.1939 | 0.1996 | 0.2903 | 0.4000 | 0.4404 | 0.4525 | 0.4603 |
|  |  | 0.3985 | 0.3366 | 0.3169 | 0.3750 | 0.4554 | 0.5122 | 0.5417 | 0.5693 |  |
| 76 | os | 0.4614 | 0.3804 | 0.2549 | 0.1401 | 0.1989 | 0.2965 | 0.3971 | 0.4521 | 0.4538 |
|  |  | 0.4310 | 0.3461 | 0.3195 | 0.3555 | 0.4355 | 0.4984 | 0.5467 | 0.5628 |  |
| 78 | PT | 0.5711 | 0.5106 | 0.3803 | 0.1694 | 0.0819 | 0.1982 | 0.3370 | 0.4127 | 0.4430 |
|  |  | 0.4431 | 0.3649 | 0.3193 | 0.3409 | 0.4138 | 0.4798 | 0.5289 | 0.5488 |  |
| 80 | H/G | 0.6538 | 0.5718 | 0.4872 | 0.3073 | 0.0806 | 0.1015 | 0.2527 | 0.3666 | 0.4265 |
|  |  | 0.4496 | 0.3826 | 0.3250 | 0.3269 | 0.3914 | 0.4466 | 0.5088 | 0.5499 |  |
| 82 | PB | 0.8557 | 0.7215 | 0.5926 | 0.3856 | 0.1469 | 0.0945 | 0.2350 | 0.3655 | 0.4384 |
|  |  | 0.4701 | 0.4067 | 0.3415 | 0.3295 | 0.3866 | 0.4381 | 0.4958 | 0.5415 |  |

Table V. Backscattering Phase Shifts $\phi_{b}(k)$ in Radian vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Herman-Skillman Wave Functions

| Z | CHEM | $\mathrm{k}=3.7795$ | 4.2519 | 4.7243 | 5.1967 | 5.6692 | 6.1416 | 6.6140 | 7.0865 | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 8.5038 | 9.4486 | 10.3935 | 11.3384 | 12.2832 | 13.2281 | 14.1729 | 15.1178 |  |
| 14 | Sl | 2.8451 | 2.7102 | 2.5684 | 2.4237 | 2.2930 | 2.1695 | 2.0314 | 1.8893 | 1.7979 |
|  |  | 1.5790 | 1.3425 | 1.1418 | 0.9229 | 0.7272 | 0.5244 | 0.3251 | 0.1337 |  |
| 17 | CL | 3.5339 | 3.4935 | 3.4518 | 3.3727 | 3.2504 | 3.1405 | 3.0158 | 2.9084 | 2.8179 |
|  |  | 2.5775 | 2.3437 | 2.1201 | 1.9182 | 1.6835 | 1.5147 | 1.3126 | 1.1424 |  |
| 20 | CA | 4.9368 | 4.9506 | 4.9006 | 4.8040 | 4.7095 | 4.5616 | 4.4313 | 4.2976 | 4.1491 |
|  |  | 3.8851 | 3.6277 | 3.3828 | 3.0205 | 2.8617 | 2.7022 | 2.5257 | 2.3615 |  |
| 40 | ZR | 5.5087 | 5.6976 | 5.9968 | 6.3248 | 6.7537 | 7.1845 | 7.3301 | 7.3862 | 7.3997 |
|  |  | 7.3259 | 7.2062 | 7.0627 | 6.8848 | 6.7291 | 6.5750 | 6.4172 | 6.2162 |  |
| 42 | MO | 5.3241 | 5.6417 | 5.9025 | 6.2277 | 6.5820 | 6.9913 | 7.3814 | 7.5168 | 7.5415 |
|  |  | 7.5779 | 7.4839 | 7.3453 | 7.1972 | 7.0307 | 6.8815 | 6.7131 | 6.5743 |  |
| 44 | RU | 5.1521 | 5.5560 | 5.8075 | 6.1610 | 6.4952 | 6.8481 | 7.2943 | 7.5823 | 7.6837 |
|  |  | 7.7764 | 7.7103 | 7.6053 | 7.4610 | 7.3223 | 7.1668 | 7.0090 | 6.8538 |  |
| 46 | PD | 4.9785 | 5.4527 | 5.7202 | 6.0619 | 6.3308 | 6.6559 | 7.1120 | 7.4861 | 7.6767 |
|  |  | 7.8956 | 7.8671 | 7.8056 | 7.6815 | 7.5570 | 7.4254 | 7.2556 | 7.1135 |  |
| 47 | AG | 4.9498 | 5.4256 | 5.7322 | 6.0104 | 6.2758 | 6.5895 | 7.0201 | 7.4073 | 7.6789 |
|  |  | 7.9364 | 7.9624 | 7.8971 | 7.7941 | 7.6706 | 7.5470 | 7.3844 | 7.2409 |  |
| 50 | SN | 6.2323 | 6.5485 | 6.7944 | 6.9256 | 7.0671 | 7.3134 | 7.5998 | 7.9268 | 8.2718 |
|  |  | 8.5524 | 8.6258 | 8.5639 | 8.4687 | 8.3487 | 8.2004 | 8.0240 | 7.8928 |  |
| 53 | 1 | 7.2328 | 7.4433 | 7.5482 | 7.6736 | 7.7480 | 7.9092 | 8.1375 | 8.3787 | 8.6663 |
|  |  | 9.0611 | 9.1503 | 9.1388 | 9.0401 | 8.9299 | 8.8010 | 8.6309 | 8.4567 |  |
| 57 | LA | 9.1054 | 9.1140 | 9.0879 | 9.0990 | 9.0917 | 9.1506 | 9.2783 | 9.4672 | 9.6694 |
|  |  | 10.0405 | 10.1739 | 10.1463 | 10.0304 | 9.9063 | 9.7864 | 9.6034 | 9.4633 |  |
| 58 | CE | 8.8314 | 8.8587 | 8.7713 | 8.8249 | 8.8322 | 8.8543 | 8.9636 | $9.1388$ | 9.3441 |
|  |  | 9.7813 | 9.9848 | 9.9920 | 9.9276 | 9.8101 | 9.7087 | 9.5561 | $9.4087$ |  |
| 65 | TB | 8.9896 | 9.0186 | 8.8981 | 8.8092 | 8.7686 | 8.7275 | 8.7431 | 8.7896 | 8.8404 |
|  |  | 9.2244 | 9.6925 | 9.9728 | 10.0507 | 10.0328 | 9.9813 | 9.9134 | 9.8120 |  |
| 70 | YB | 9.1561 | 9.1654 | 9.1020 | 8.9196 | 8.8057 | 8.7093 | 8.6750 | 8.6419 | 8.6522 |
|  |  | 8.8262 | 9.2384 | 9.7232 | 9.9597 | 10.0334 | 10.0505 | 10.0326 | 9.9615 |  |
| 74 | W | 10.8847 | 10.8005 | 10.3731 | 9.7269 | 9.3337 | 9.2044 | 9.1181 | 9.0581 | 9.0780 |
|  |  | 9.1798 | 9.5185 | 9.9579 | 10.2952 | 10.4195 | 10.4622 | 10.4376 | 10.3847 |  |
| 76 | OS | $10.9741$ | $11.2710$ | $11.3216$ | $10,2771$ | 9.5414 | 9.2519 | 9.1696 | $9.1667$ | 9.1447 |
|  |  | 9.2602 | 9.5576 | 9.9875 | 10.3635 | 10.5346 | 10.6037 | 10.5956 | 10.5386 |  |
| 78 | PT | $11.1468$ | 11.5015 | 11.6393 | 11.5297 | 10.1376 | 9.3255 | $9.2371$ | $9.2168$ | 9.1978 |
|  |  | 9.3455 | 9.6205 | 10.0458 | 10.4390 | 10.6673 | 10.7535 | 10.7584 | 10.7178 |  |
| 80 | HG | 11.5236 | 11.6950 | 11.8774 | 12.0838 | 11.8658 | 9.2967 | 9.2339 | 9.3059 | 9.3491 |
|  |  | 9.4588 | 9.7259 | 10.1046 | 10.4877 | 10.7982 | 10.9296 | 10.9448 | 10.9055 |  |
| 82 | PB | 12.3853 | 12.5708 | 12.7450 | 12.9479 | 13.2543 | 8.9454 | 9.4699 | 9.6543 | 9.7369 |
|  |  | 9.8628 | 10.1082 | 10.4315 | 10.8212 | 11.1250 | 11.2779 | 11.2993 | 11,2529 |  |

states, though small but significant, can be compensated by changing the threshold energy (vide infra). The valence electronic configuration for the three transition metal series were $3 \mathrm{~d}^{z-204} 4 \mathrm{~s}^{2}, 4 \mathrm{~d}^{z-37} 5 \mathrm{~s}^{1}$, and $5 \mathrm{~d}^{z-70} 6 \mathrm{~s}^{2}$. Again, the difference between various electronic configurations can largely be compensated by $E_{0}$ variation. We also should point out that the Herman-Skillman wave functions are inadequate for heavy atoms (beyond the rare earths, for instance) because of relativistic corrections. We have used relativistic wave functions for tungsten and the result is shown in a later section. We find, however, that the difference is sufficiently small to justify the use of the more readily available nonrelativistic HermanSkillman wave functions for all elements.

## Results and Discussion

In this paper, we report ab initio theoretical EXAFS functions in the range of $k \simeq 4-15 \AA^{-1}$. The truncation at low $k$ value ( $k \simeq 4 \AA^{-1}$ ) is due to the fact that the theory is less reliable for $k \leq 4 \AA^{-1}$ as a result of inadequate treatment of valence electrons, particularly for light atoms with $Z \leqslant 9$ where the energy of the valence electrons is a substantial portion of that of the core electrons. Furthermore, other physical phenomena such as multiple scattering may become important at low $k$ values. At high $k$ values ( $k \gtrsim 15 \AA^{-1}$ ), the EXAFS signal is generally attenuated substantially by Debye-Waller factor.

Throughout this paper, we chose to use the "experimental" unit of angstroms for the amplitude function $F(k)$ (cf. Tables 1 and IV), radian for the phase functions $\phi(k)$ (cf. Tables II, $111, \mathrm{~V}-\mathrm{V} 11 \mathrm{I}$ ), and $\AA^{-1}$ for the electron wave vector $k$. It should also be noted that the functions $F(k), \phi_{b}(k)$, and $\phi_{a}{ }^{\prime}(k)$ reported in this paper are equivalent to the functions $f(\pi, k)$, $\theta(k)$, and $2 \delta_{\prime}^{\prime}(k)$ (in atomic units) used in Lee and Beni's paper. ${ }^{8}$ Furthermore, if we have an absorbing atom $A$ and a backscattering atom B , we denote the phase function by $\phi_{a b}$ which is given by
$\phi_{a b}(k)=\phi_{a}{ }^{\prime}(k)+\phi_{b}(k)-\pi, l=I$ for $K$ and $L_{1}$ edges

$$
\begin{equation*}
\phi_{a b}(k)=\phi_{a}^{\prime}(k)+\phi_{b}(k), l=2,0 \text { for } L_{11,111} \text { edges } \tag{6a}
\end{equation*}
$$

where $\phi_{a}{ }^{\prime}$ is the phase shift of the central atom A and $\phi_{b}$ is the phase shift of the neighboring atom $B$.

A detailed comparison of Tables 1-111 (Clementi-Roetti ${ }^{26 a}$ wave functions) with Tables IV-V1II (Herman-Skillman ${ }^{26 \mathrm{~b}}$ wave functions) revealed that the two sets of results agree quite well with the exception of central atom (absorber) phase shifts. A comparison of Table 111 with the corresponding Table VII showed that the $(Z+1)$ approximation in the former case results in more positive phase shifts. The difference (ca. 0.7-0.2 rad ), however, decreases with increasing $k$ values and therefore can be substantially removed by changing the energy threshold

Table VI. Central Atom Phase Shifts $\phi_{a}{ }^{l=0}(k)$ in Radian vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Herman-Skillman Wave Functions

| Z | CHEM | $\begin{array}{r} \mathrm{k}= \\ \begin{array}{r} 3.7795 \\ 8.5038 \end{array} \end{array}$ | $\begin{aligned} & 4.2519 \\ & 9.4486 \end{aligned}$ | $\begin{array}{r} 4.7243 \\ 10.3935 \end{array}$ | $\begin{array}{r} 5.1967 \\ 11.3384 \end{array}$ | $\begin{array}{r} 5.6692 \\ 12.2832 \end{array}$ | $\begin{array}{r} 6.1416 \\ 13.2281 \end{array}$ | $\begin{array}{r} 6.6140 \\ 14.1729 \end{array}$ | $\begin{array}{r} 7.0865 \\ 15.1178 \end{array}$ | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 11 | $\mathrm{NA}+$ | -5.6231 | -6.3933 | -6.8848 | -7.3767 | -7.8876 | -8.2287 | -8.6431 | -8.9622 | -9.2652 |
|  |  | -9.8182 | -10.3291 | -10.7726 | -11.1661 | -11.5138 | -11.8216 | -12.1021 | -12.3581 |  |
| 12 | $\mathrm{MG}++$ | -3.8448 | -4.6324 | -5.3798 | -5.8169 | -6.3347 | -6.8851 | -7.2518 | -7.6238 | -8.0287 |
|  |  | -8.6331 | -9.1791 | -9.7016 | -10.1483 | -10.5207 | -10.8800 | -11.2092 | -11.5071 |  |
| 14 | S1 | -5.1257 | -5.7878 | -6.3136 | -6.7332 | -7.1618 | -7.5758 | -7.9339 | -8.2163 | -8.5254 |
|  |  | -9.0594 | -9.5388 | -9.9391 | -10.3355 | -10.6638 | -10.9752 | -11.2682 | -11.5251 |  |
| 17 | CL | -3.9386 | -4.5306 | -5.0641 | -5.5371 | -5.9210 | -6.3182 | -6.7155 | -7.0543 | -7.3263 |
|  |  | -7.9082 | -8.3831 | -8.8327 | -9.2170 | -9.5822 | -9.9003 | -10.2143 | -10.4905 |  |
| 19 | K + | -0.8143 | -1.6909 | -2.2709 | -2.9487 | -3.4504 | -3.9979 | -4.4169 | -4.8718 | -5.2530 |
|  |  | -5.9536 | -6.5740 | -7.1175 | -7.5960 | -8.0316 | -8.4189 | .8.7911 | -9.1260 |  |
| 20 | $\mathrm{CA}++$ | 0.6406 | -0.0014 | -0.8699 | -1.4697 | -2.0743 | -2.6683 | -3.1234 | -3.6604 | -4.0268 |
|  |  | -4.8218 | -5.5110 | -6.1175 | -6.6567 | -7.1396 | -7.5766 | -7.9730 | .8.3377 |  |
| 20 | CA | -1.6038 | -2.3361 | -2.8941 | -3.4800 | -3.9310 | -4.4009 | -4.8048 | -5.1989 | -5.5518 |
|  |  | -6.2042 | -6.7698 | -7.2768 | -7.7304 | -8.1389 | -8.5121 | -8.8554 | -9.1691 |  |
| 22 | Tl | -1.3496 | -2.0813 | -2.6117 | -3.1165 | -3.6305 | -4.0214 | -4.3987 | -4.8077 | .5.1324 |
|  |  | -5.7680 | -6.3163 | -6.8133 | -7.2728 | -7.6814 | -8.0480 | -8.3895 | -8.7136 |  |
| 26 | FE | -0.5981 | -1.2789 | -1.8973 | -2.3640 | -2.7732 | -3.2195 | -3.6079 | -3.9363 | -4.2863 |
|  |  | -4.9006 | -5.4475 | -5.9420 | -6.3928 | -6.7960 | -7.1690 | -7.5194 | .7.8358 |  |
| 28 | N1 | -0.2909 | -0.9459 | -1.5472 | -2.0204 | -2.4204 | -2.8508 | -3.2421 | -3.5710 | -3.9074 |
|  |  | -4.5158 | -5.0653 | -5.5374 | -5.9978 | -6.4019 | -6.7785 | -7.1283 | -7.4465 |  |
| 32 | GE | 0.8703 | 0.2092 | -0.2978 | -0.8170 | -1.3128 | -1.7093 | -2.0806 | -2.4920 | -2.8277 |
|  |  | -3.4527 | -4,0051 | -4.5191 | -4.9725 | -5.3879 | -5.7904 | -6.1490 | -6.4830 |  |
| 35 | BR | 1.7659 | 1.1103 | 0.5794 | 0.0977 | -0.4215 | -0.8458 | -1.2233 | -1.6066 | -1.9886 |
|  |  | -2.6082 | -3,2051 | -3.7131 | -4.1773 | -4.6187 | -5.0200 | -5.3906 | -5.7394 |  |
| 40 | ZR | 3.8309 | 3.1251 | 2.5669 | 1.9563 | 1.4417 | 1.0104 | 0.5316 | 0.1307 | -0.2532 |
|  |  | -0.9471 | -1.5886 | -2.1614 | -2.6682 | -3.1331 | -3.5692 | -3.9683 | .4.3512 |  |
| 42 | MO | 4.2333 | 3.4817 | 2.8584 | 2,3604 | 1.8328 | 1.3441 | 0.9345 | 0.5199 | 0.1211 |
|  |  | -0.5622 | -1.1870 | -1.7663 | -2.2750 | -2.7419 | -3.1818 | -3.5866 | -3.9524 |  |
| 44 | RU | 4.5455 | 3.8317 | 3.1836 | 2.6907 | 2.1956 | 1.6875 | 1.2868 | 0.9049 | 0.5003 |
|  |  | -0.1749 | -0.8243 | -1.3863 | -1.8955 | -2.3768 | -2.8095 | -3.2053 | -3.5863 |  |
| 46 | PD | 4.8624 | 4.1300 | 3.5221 | 3.0478 | 2.5299 | 2.0388 | 1.6580 | 1.2552 | 0.8468 |
|  |  | 0.1802 | -0.4482 | -1.0313 | -1.5420 | -2.0109 | -2.4527 | -2.8626 | -3.2327 |  |
| 50 | SN | 6.1712 | 5.5288 | 4.8544 | 4.2728 | 3.8081 | 3.2835 | 2.8175 | 2.4217 | 1.9980 |
|  |  | 1.3042 | 0.6382 | 0.0332 | -0.5017 | -0.9909 | -1.4469 | -1.8849 | -2.2776 |  |
| 53 | 1 | 6.9861 | 6.3488 | 5.6996 | 5.0999 | 4.5908 | 4.0966 | 3.6020 | 3.1848 | 2.7917 |
|  |  | 2.0548 | 1.3868 | 0.7883 | 0.2369 | -0.2833 | -0.7488 | -1.1859 | -1.5902 |  |
| 57 | LA | 8.9217 | 8.2145 | 7.4576 | 6.8800 | 6.2577 | 5.7083 | 5.2118 | 4.7188 | 4.2983 |
|  |  | 3.5052 | 2.7852 | 2.1321 | 1.5268 | 0.9808 | 0.4739 | -0.0014 | -0.4375 |  |
| 58 | CE | 8.8166 | 8.1511 | 7.4003 | 6.8131 | 6.2281 | 5.6672 | 5.1862 | 4.6962 | 4.2806 |
|  |  | 3.4880 | 2.7814 | 2.1388 | 1.5480 | 1.0024 | 0.5109 | 0.0419 | -0.4007 |  |
| 64 | $G D$ | 9.4593 | 8.8153 | 8.0828 | 7.4980 | 6.9600 | 6.3999 | 5.9272 | 5.4477 | 5.0296 |
|  |  | 4.2472 | 3.5518 | 2.9159 | 2.3326 | 1.8002 | 1.3049 | 0.8393 | 0.4053 |  |
| 70 | YB | 9.7131 | 9.0941 | 8.4302 | 7.8181 | 7.2945 | 6.7735 | 6.3181 | 5.8661 | 5.4388 |
|  |  | 4.6757 | 4.0225 | 3.4051 | 2.8361 | 2.3099 | 1,8207 | 1.3664 | 0.9411 |  |
| 74 | W | 10.4268 | 9.6477 | 9.0202 | 8.4815 | 7.9303 | 7.3854 | 6.9506 | 6.5046 | 6.0647 |
|  |  | 5.3324 | 4.6571 | 4.0207 | 3.4588 | 2.9322 | 2.4358 | 1.9818 | 1.5598 |  |
| 76 | OS | 10.6930 | 9.9356 | 9.2906 | 8.7545 | 8.2274 | 7.6945 | 7.2307 | 6.8067 | 6.3706 |
|  |  | 5.6367 | 4.9411 | 4.3220 | 3.7535 | 3.2160 | 2.7260 | 2.2752 | 1.8352 |  |
| 78 | PT | 10.9689 | 10.2018 | 9.5890 | 9.0822 | 8.5108 | 7.9961 | 7.5602 | 7.0946 | 6.6577 |
|  |  | 5.9321 | 5.2571 | 4.6173 | 4.0381 | 3.5196 | 3.0234 | 2.5544 | 2.1299 |  |
| 80 | HIG | 11.2355 | 10.5896 | 9.9143 | 9.3646 | 8.8521 | 8.2996 | 7.8570 | 7.4171 | 6.9807 |
|  |  | 6.2308 | 5.5580 | 4.9246 | 4.3319 | 3.8022 | 3.3126 | 2.8496 | 2.4138 |  |
| 82 | PB | 11.9467 | 11.2913 | 10.5835 | 10.0159 | 9.4675 | 8.9075 | 8.4542 | 7.9969 | 7.5528 |
|  |  | 6.7782 | 6.0730 | 5.4323 | 4.8243 | 4.2836 | 3.7834 | 3.3072 | 2.8597 |  |

(vide infra). Furthermore, it should be noted that the phase functions are listed and plotted as an increasing function of atomic number $Z$ merely for clarity and for the purpose of facilitating interpolation of the phase shifts of intermediate $Z$ elements. In practice, any phase functions can be modified by $\pm 2 n \pi$ where $n=0,1,2,3, \ldots$, since $\sin (\phi(k) \pm 2 n \pi)=$ $\sin (\phi(k))$. With these remarks, we shall now discuss the backscattering amplitude $(F(k)$ ), the backscattering phase $\left(\phi_{b}(k)\right)$, and the central atom phase shifts $\left(\phi_{a}(k)\right)$ with the aid of Figures 2-14.

Amplitude. Figures 2 a and 2 b depict the scattering amplitudes for two main groups 7A and 4A, respectively, whereas Figure 2 c shows the amplitude functions of the transition
metals $\mathrm{Fe}, \mathrm{Ru}$, and Os . It is apparent that as the atomic number $Z$ increases the scattering amplitude at high $k$ values generally increases. More importantly, there are peaks and valleys in the amplitude functions which move to higher $k$ values as $Z$ increases. These amplitude peaks correspond to the resonances of the electron-atom scattering process and can be associated with the $1 / 2 \pi$ crossings of phase shifts with different $l$ values. Figures 3a-d show how these peaks and valleys progress within each series. For light atoms with $Z \leqq 10$, the amplitude function peaks at the low $k$ region ( $\leqslant 3 \AA^{-1}$ ) such that only a monotonically decreasing function is observed. This structureless tail is due to the fact that, when the electron exceeds the binding energy of the deepest shell, the electron is

Table VII. Central Atom Phase Shifts $\phi_{a}{ }^{\prime=1}(k)$ in Radian vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Herman-Skillman Wave Functions

| Z | CHEM | $\begin{array}{r} \mathrm{k}=3.7795 \\ 8.5038 \end{array}$ | $\begin{aligned} & 4.2519 \\ & 9.4486 \end{aligned}$ | $\begin{array}{r} 4.7243 \\ 10.3935 \end{array}$ | $\begin{array}{r} 5.1967 \\ 11.3384 \end{array}$ | $\begin{array}{r} 5.6692 \\ 12.2832 \end{array}$ | $\begin{array}{r} 6.1416 \\ 13.2281 \end{array}$ | $\begin{array}{r} 6.6140 \\ 14.1729 \end{array}$ | $\begin{array}{r} 7.0865 \\ 15.1178 \end{array}$ | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 6 | C | .7.0744 | -7.2817 | -7.4487 | .7.6994 | -7.9676 | -8.1821 | -8.3345 | -8.4618 | -8.6018 |
|  |  | -8.8888 | -9.0707 | -9.2798 | -9.4242 | -9.5691 | -9.6967 | -9.8045 | .9.9151 |  |
| 8 | 0 | -6.5189 | -6.7480 | -6.9315 | -7.1700 | -7.4127 | -7.6278 | -7.7959 | .7.9196 | -8.0470 |
|  |  | -8.3582 | -8.5693 | -8.7617 | -8.9497 | -9.0839 | -9.2360 | -9.3545 | -9.4670 |  |
| 9 | F | -6.2972 | -6.5199 | -6.7071 | -6.9492 | -7.2004 | -7.4174 | -7.5870 | -7.7204 | -7.8349 |
|  |  | -8.1329 | -8.3485 | -8.5462 | -8.7425 | -8.8804 | -9.0351 | -9.1566 | -9.2720 |  |
| 11 | $\mathrm{NA}+$ | -3.2155 | -3.8481 | -4.2669 | -4.6247 | -5.0507 | -5.2996 | -5.6235 | -5.8562 | -6.0813 |
|  |  | -6.4784 | -6.8412 | -7.1488 | -7.4189 | -7.6600 | -7.8716 | -8.0583 | -8.2261 |  |
| 12 | $\mathrm{MG}++$ | -1.4390 | -2.0405 | -2.7238 | -3.0884 | -3.4938 | -3.9705 | -4.2603 | -4.5530 | -4.8926 |
|  |  | -5.3568 | -5.7625 | -6.1516 | -6.4918 | -6.7651 | -7.0197 | -7.2475 | -7.4645 |  |
| 14 | S 1 | -2.8651 | -3.4002 | -3.8494 | -4.1870 | -4.5117 | -4.8393 | -5.1257 | -5.3292 | .5.5617 |
|  |  | -5.9614 | -6.3161 | -6.5963 | -6.8764 | -7.1045 | -7.3137 | -7.5183 | -7.6926 |  |
| 17 | CL | -1.7939 | -2.2287 | -2.6765 | -3.0897 | -3.3912 | -3.6954 | -4.0199 | -4.2977 | -4.5038 |
|  |  | -4.9602 | -5.3250 | -5.6685 | -5.9513 | -6.2241 | -6.4524 | -6.6771 | -6.8757 |  |
| 19 | $K+$ | 1.3571 | 0.6147 | 0.0905 | -0.4883 | -0.9257 | -1.4047 | -1.7551 | -2.1531 | $-2.4673$ |
|  |  | -3.0579 | -3.5781 | -4.0250 | -4.4108 | . 4.7602 | -5.0656 | -5.3588 | -5.6179 |  |
| 20 | $\mathrm{CA}++$ | 2.9048 | 2.2965 | 1.5766 | 0.9778 | 0.4902 | -0.0731 | -0.4507 | -0.9338 | -1.2464 |
|  |  | -1.9394 | -2.5310 | -3.0439 | -3.4957 | -3.8939 | -4.2513 | -4.5735 | -4.8644 |  |
| 20 | CA | 0.5733 | -0.0952 | -0.5599 | -1.0552 | -1.4490 | -1.8348 | -2.1860 | -2.5083 | $-2.8093$ |
|  |  | -3.3537 | -3.8148 | -4.2257 | -4.5900 | -4.9121 | -5.2042 | -5.4720 | -5.7105 |  |
| 22 | TI | 0.7928 | 0.1399 | -0.3505 | -0.7510 | -1.1891 | . 1.5297 | -1.8307 | -2.1774 | -2.4535 |
|  |  | -2.9774 | -3.4271 | -3.8348 | -4.2071 | -4.5351 | -4.8252 | -5.0924 | -5.3461 |  |
| 26 | FE | 1.3837 | 0.8536 | 0.3077 | -0.1080 | -0.4466 | -0.8114 | -1.1523 | -1.4268 | -1.7096 |
|  |  | -2.2304 | -2.6766 | -3.0871 | -3.4565 | -3.7820 | -4.0841 | -4.3658 | -4.6186 |  |
| 28 | N1 | 1.6350 | 1.1435 | 0.6110 | 0.1706 | -0.1425 | -0.4901 | -0.8352 | -1.1129 | -1.3810 |
|  |  | -1.8995 | -2.3468 | -2.7365 | -3.1171 | -3.4458 | -3.7537 | -4.0366 | -4.2916 |  |
| 32 | GE | 2.8531 | 2.2248 | 1.7630 | 1.3573 | 0.9298 | 0.5690 | 0.2700 | -0.0767 | -0.3758 |
|  |  | -0.8942 | -1.3665 | -1.7991 | -2.1741 | -2.5210 | -2.8564 | -3.1522 | -3.4270 |  |
| 35 | BR | 3.7269 | 3.1353 | 2.6303 | 2.2422 | 1.8157 | 1.4317 | 1.0971 | 0.7827 | 0.4517 |
|  |  | -0.0860 | -0.5993 | -1.0274 | -1.4246 | -1.7951 | -2.1336 | -2.4457 | -2.7370 |  |
| 40 | ZR | 5.8079 | 5.1256 | 4.6373 | 4.1281 | 3.6490 | 3.2773 | 2.8561 | 2.4865 | 2.1632 |
|  |  | 1.5447 | 0.9799 | 0.4818 | 0.0414 | -0.3642 | -0.7440 | -1.0869 | -1.4163 |  |
| 42 | MO | 6.1471 | 5.5279 | 4.9250 | 4.4785 | 4.0464 | 3.6009 | 3.2270 | 2.8816 | 2.5219 |
|  |  | 1.9237 | 1.3660 | 0.8583 | 0.4160 | 0.0058 | -0.3779 | -0.7269 | -1.0414 |  |
| 44 | RU | 6.4206 | 5.8617 | 5.2665 | 4.7926 | 4.3806 | 3.9385 | 3.5686 | 3.2431 | 2.8949 |
|  |  | 2.2888 | 1.7143 | 1.2250 | 0.7741 | 0.3513 | -0.0242 | -0.3677 | -0.6993 |  |
| 46 | PD | 6.7635 | 6.1581 | 5.5692 | 5.1463 | 4.7175 | 4.2611 | 3.9233 | 3.5903 | 3.2188 |
|  |  | 2.6376 | 2.0726 | 1.5596 | 1.1130 | 0.6981 | 0.3114 | -0.0450 | -0.3665 |  |
| 47 | $A G$ | 7.4130 | 7.0555 | 6.6507 | 6.3647 | 6.0790 | 5.7689 | 5.5393 | 5.3076 | 5.0642 |
|  |  | 4.6584 | 4.2777 | 3.9214 | 3.6012 | 3.3222 | 3.0718 | 2.8187 | 2.6021 |  |
| 50 | SN | 8.0817 | 7.4989 | 6.9398 | 6.3880 | 5.9590 | 5.5205 | 5.0864 | 4.7366 | 4.3678 |
|  |  | 3.7451 | 3.1433 | 2.6045 | 2.1309 | 1.6957 | 1.2903 | 0.9009 | 0.5570 |  |
| 53 | 1 | 8.9247 | 8.3161 | 7.7791 | 7.2392 | 6.7528 | 6.3289 | 5.8870 | 5.4941 | 5.1532 |
|  |  | 4.4825 | 3.8899 | 3.3529 | 2.8544 | 2.3887 | 1.9751 | 1.5843 | 1.2241 |  |
| 55 | CS + | 10.3098 | 9.6092 | 8.8908 | 8.3605 | 7.9268 | 7.4458 | 6.9389 | 6.5101 | 6.1637 |
|  |  | 5.4374 | 4.8397 | 4.2264 | 3.7234 | 3.2053 | 2.7855 | 2.3499 | 1.9675 |  |
| 57 | LA | 10.8567 | 10.2554 | 9.5608 | 9.0124 | 8.4708 | 7.9414 | 7.5065 | 7.0439 | 6.6639 |
|  |  | 5.9371 | 5.2814 | 4.6881 | 4.1370 | 3.6430 | 3.1842 | 2.7531 | 2.3596 |  |
| 58 | CE | 10.7507 | 10.1561 | 9.4991 | 8.9259 | 8.4226 | 7.8903 | 7.4572 | 7.0182 | 6.6274 |
|  |  | 5.9071 | 5.2682 | 4.6874 | 4.1510 | 3.6564 | 3.2123 | 2.7865 | 2.3858 |  |
| 64 | GD | 11.3615 | 10.7710 | 10.1419 | 9.5739 | 9.1174 | 8.6007 | 8.1617 | 7.7395 | 7.3430 |
|  |  | 6.6353 | 6.0082 | 5.4315 | 4.9007 | 4.4168 | 3.9662 | 3.5430 | 3.1502 |  |
| 70 | $Y B$ | 11.5744 | 10.9738 | 10.4332 | 9.8476 | 9.3850 | 8.9363 | 8.5022 | 8.1120 | 7.7131 |
|  |  | 7.0261 | 6.4381 | 5.8762 | 5.3578 | 4.8803 | 4.4379 | 4.0281 | 3.6442 |  |
| 74 | W | 12.2519 | 11.6107 | 10.9990 | 10.4924 | 10.0398 | 9.5418 | 9.1289 | 8.7431 | 8.3425 |
|  |  | 7.6748 | 7.0550 | 6.4830 | 5.9731 | 5.4897 | 5.0400 | 4.6301 | 4.2456 |  |
| 76 | OS | 12,5102 | 11.8931 | 11.2782 | 10.7721 | 10.3217 | 9.8559 | 9.4097 | 9.0322 | 8.6483 |
|  |  | 7.9683 | 7.3377 | 6.7810 | 6.2573 | 5.7681 | 5.3254 | 4.9137 | 4.5123 |  |
| 78 | PT | 12.8279 | 12.1642 | 11.5505 | 11.0983 | 10.6215 | 10.1311 | 9.7348 | 9.3312 | 8.9200 |
|  |  | 8.2696 | 7.6463 | 7.0656 | 6.5410 | 6.0655 | 5.6126 | 5.1879 | 4.8013 |  |
| 80 | HG | 13.0831 | 12.4816 | 11.9145 | 11.3844 | 10.9332 | 10.4476 | 10.0268 | 9.6428 | 9.2403 |
|  |  | 8.5622 | 7.9474 | 7.3650 | 6.8229 | 6.3431 | 5.8984 | 5.4754 | 5.0767 |  |
| 82 | PB | 13.7859 | 13.2106 | 12.6005 | 12.0451 | 11.5681 | 11.0660 | 10.6298 | 10.2276 | 9.8130 |
|  |  | 9.1126 | 8.4636 | 7.8730 | 7.3162 | 6.8247 | 6.3675 | 5.9295 | 5.5188 |  |

sampling mostly the nuclear potential. For elements with 10 $\leqq Z \leqq 30$, we observe an amplitude envelope whose peak height decreases while its peak position advances in $k$ as $Z$ increases. For elements with $30 \leqq Z \leqq 54$, we find one peak
and one valley in the amplitude function. A second peak starts to come in from the low $k$ region. Both peaks increase in amplitude and advance in peak position as $Z$ increases. For elements with $57 \leqq Z \leqq 71$, both peaks advance in peak position

Table VIII. Central Atom Phase Shifts $\phi_{a}{ }^{l=2}(k)$ in Radian vs. Photoelectron Wave Vector $k$ in $\AA^{-1}$, Calculated Using Herman-Skillman Wave Functions

| Z | CHEM | $k=3.7795$ | 4.2519 | 4.7243 | 5.1967 | 5.6692 | 6.1416 | 6.6140 | 7.0865 | 7.5589 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 8.5038 | 9.4486 | 10.3935 | 11.3384 | 12.2832 | 13.2281 | 14.1729 | 15.1178 |  |
| 11 | NA+ | -0.3995 | -0.8132 | -0.9636 | -1.1910 | -1.4274 | -1.5356 | -1.7250 | -1.8052 | -1.9704 |
|  |  | -2.1820 | -2.3796 | -2.5390 | -2.6900 | -2.8270 | -2.9523 | -3.0639 | -3.1649 |  |
| 12 | $\mathrm{MG}++$ | 1.3366 | 0.9309 | 0.4564 | 0.3056 | -0.0218 | -0.3142 | -0.4408 | -0.6764 | -0.8554 |
|  |  | -1.1806 | -1.4292 | -1.6413 | -1.8642 | -2.0579 | -2.2275 | -2.3535 | -2.4818 |  |
| 14 | S1 | -0.7863 | -1.0065 | -1.2323 | -1.3593 | -1.4973 | -1.6609 | -1.7749 | -1.8344 | -1.9419 |
|  |  | -2.1014 | -2.2746 | -2.4059 | -2.5426 | -2.6554 | -2.7539 | -2.8504 | -2.9472 |  |
| 17 | CL | -0.2270 | -0.2936 | -0.4897 | -0.7129 | -0.8225 | -0.9409 | -1.1002 | -1.2162 | -1.2796 |
|  |  | -1.5121 | -1.6571 | -1.8275 | -1.9597 | -2.0875 | -2.2033 | -2.3084 | -2.4182 |  |
| 19 | K + | 2.9156 | 2.5286 | 2.1377 | 1.8014 | 1.5062 | 1.2084 | 1.0116 | 0.7518 | 0.5703 |
|  |  | 0.2063 | -0.1042 | -0.3633 | -0.5900 | -0.8017 | -0.9859 | -1.1670 | -1.3169 |  |
| 20 | $\mathrm{CA}++$ | 4.7443 | 4.1594 | 3.7548 | 3.2475 | 2.9594 | 2.5239 | 2.2975 | 1.9436 | 1.7547 |
|  |  | 1.2801 | 0.8908 | 0.5610 | 0.2602 | 0.0020 | -0.2311 | -0.4373 | -0.6244 |  |
| 20 | CA | 2.0448 | 1.6636 | 1.3692 | 1.1028 | 0.8626 | 0.6594 | 0.4565 | 0.2822 | 0.1137 |
|  |  | -0.1949 | -0.4470 | -0.6673 | -0.8742 | -1.0508 | -1.2141 | -1.3598 | -1.4922 |  |
| 22 | Tl | 2.0353 | 1.7127 | 1.3722 | 1.1717 | 0.9422 | 0.7419 | 0.6008 | 0.4051 | 0.2559 |
|  |  | -0.0259 | -0.2628 | -0.4696 | -0.6788 | -0.8663 | -1.0169 | -1.1526 | -1.2932 |  |
| 26 | FE | 2.2020 | 1.9926 | 1.7277 | 1.4388 | 1.2717 | 1.1124 | 0.9207 | 0.7721 | 0.6362 |
|  |  | 0.3515 | 0.1303 | -0.0759 | -0.2748 | -0.4459 | -0.6032 | -0.7529 | -0.8865 |  |
| 28 | N1 | 2.2964 | 2.0904 | 1.8684 | 1.5675 | 1.4034 | 1.2655 | 1.0768 | 0.9185 | 0.8013 |
|  |  | 0.5233 | 0.3003 | 0.1068 | -0.0928 | -0.2627 | -0.4257 | -0.5750 | -0.7118 |  |
| 32 | GE | 3.4808 | 3.1522 | 2.7816 | 2.5499 | 2.3530 | 2.1165 | 1.9519 | 1.7749 | 1.5910 |
|  |  | 1.3132 | 1.0420 | 0.8017 | 0.6026 | 0.4131 | 0.2241 | 0.0594 | -0.0915 |  |
| 35 | BR | 4.2896 | 4.0442 | 3.6549 | 3.3775 | 3.1748 | 2.9537 | 2.7140 | 2.5398 | 2.3480 |
|  |  | 2.0083 | 1.7061 | 1.4580 | 1.2218 | 1.0069 | 0.8098 | 0.6283 | 0.4587 |  |
| 40 | ZR | 6.3946 | 6.0007 | 5.6098 | 5.3205 | 4.9774 | 4.7086 | 4.4491 | 4.1646 | 3.9606 |
|  |  | 3.5365 | 3.1589 | 2.8311 | 2.5430 | 2.2758 | 2.0247 | 1.8051 | 1.5931 |  |
| 42 | MO | 6.6133 | 6.3255 | 5.9501 | 5.5844 | 5.3187 | 5.0429 | 4.7477 | 4.5307 | 4.2950 |
|  |  | 3.8789 | 3.4912 | 3.1549 | 2.8612 | 2.5850 | 2.3293 | 2.1041 | 1.9009 |  |
| 44 | RU | 6.8945 | 6.5710 | 6.2631 | 5.8888 | 5.5835 | 5.3282 | 5.0698 | 4.8339 | 4.6216 |
|  |  | 4.1896 | 3.8014 | 3.4727 | 3.1599 | 2.8769 | 2.6276 | 2.4024 | 2.1809 |  |
| 46 | PD | 7.1714 | 6.8840 | 6.5244 | 6.1577 | 5.9111 | 5.6099 | 5.3509 | 5.1543 | 4.9069 |
|  |  | 4.5062 | 4.1031 | 3.7560 | 3.4523 | 3.1646 | 2.9073 | 2.6681 | 2.4569 |  |
| 50 | SN | 8.7603 | 8.2233 | 7.8591 | 7.4918 | 7.1216 | 6.8426 | 6.5317 | 6.2581 | 6.0202 |
|  |  | 5.5650 | 5.1250 | 4.7401 | 4.4031 | 4.0928 | 3.8059 | 3.5254 | 3.2909 |  |
| 53 | 1 | 9.6422 | 9.1056 | 8.6831 | 8.3551 | 7.9584 | 7.6306 | 7.3421 | 7.0258 | 6.7741 |
|  |  | 6.2836 | 5.8591 | 5.4556 | 5.0838 | 4.7453 | 4.4476 | 4.1595 | 3.9002 |  |
| 57 | LA | 11.6200 | 11.0909 | 10.6303 | 10.1350 | 9.7425 | 9.3073 | 8.9637 | 8.6106 | 8.2968 |
|  |  | 7.7315 | 7.2254 | 6.7693 | 6.3422 | 5.9646 | 5.6144 | 5.2813 | 4.9833 |  |
| 58 | CE | 11.4894 | 10.9460 | 10.5122 | 10.0288 | 9.6359 | 9.2402 | 8.8656 | 8.5562 | 8.2289 |
|  |  | 7.6787 | 7.1910 | 6.7468 | 6.3326 | 5.9523 | 5.6157 | 5.2869 | 4.9811 |  |
| 64 | GD | 11.9957 | 11.4600 | 11.0329 | 10.5967 | 10.2245 | 9.8649 | 9.4819 | 9.1841 | 8.8624 |
|  |  | 8.3220 | 7.8415 | 7.3944 | 6.9822 | 6.6076 | 6.2605 | 5.9345 | 5.6351 |  |
| 70 | YB | 12.0252 | 11.5305 | 11.1428 | 10.7483 | 10.3436 | 10.0665 | 9.7133 | 9.4184 | 9.1284 |
|  |  | 8.5982 | 8.1509 | 7.7170 | 7.3189 | 6.9550 | 6.6196 | 6.3096 | 6.0174 |  |
| 74 | W | 12.5569 | 12.1504 | 11.7813 | 11.3239 | 10.9935 | 10.6783 | 10.3303 | 10.0274 | 9.7497 |
|  |  | 9.2114 | 8.7382 | 8.3088 | 7.9078 | 7.5280 | 7.1859 | 6.8713 | 6.5722 |  |
| 76 | OS | 12.8359 | 12.4158 | 12.0424 | 11.6109 | 11.2547 | 10.9740 | 10.6219 | 10.3022 | 10.0360 |
|  |  | 9.4885 | 9.0202 | 8.5885 | 8.1705 | 7.7954 | 7.4539 | 7.1305 | 6.8183 |  |
| 78 | PT | 13.1248 | 12.7441 | 12.3083 | 11.8942 | 11.5823 | 11.2487 | 10.9034 | 10.6135 | 10.3072 |
|  |  | 9.7937 | 9.3002 | 8.8599 | 8.4476 | 8.0702 | 7.7199 | 7.3916 | 7.0878 |  |
| 80 | HIG | 13.5611 | 13.0062 | 12.5998 | 12.2395 | 11.8503 | 11.5302 | 11.2065 | 10.9026 | 10.6126 |
|  |  | 10.0773 | 9.5912 | 9.1331 | 8.7092 | 8.3387 | 7.9918 | 7.6568 | 7.3423 |  |
| 82 | PB | 14.3165 | 13.7561 | 13.3507 | 12.9237 | 12.5148 | 12.1799 | 11.8216 | 11.5009 | 11.1963 |
|  |  | 10.6319 | 10.1103 | 9.6426 | 9.2027 | 8.8174 | 8.4533 | 8.1012 | 7.7740 |  |

(in $k$ ) while diminish somewhat in peak height. Finally, for elements with $72 \leqslant Z \leqslant 82$, a third peak starts to come in from the low $k$ region. The three peaks and the two valleys move to higher $k$ values as $Z$ increases.

The positions of the peaks and valleys of the amplitude functions of some representative elements are plotted as a function of atomic number $Z$ in Figure 8. It is readily apparent that roughly linear relationships exist for each peak and valley. Periodic deviations from such linear relationships are probably due to variation in electronic configurations. However, such deviations are not very significant in comparison to the present accuracy of EXAFS data. Nevertheless, the simple linear
trends of these peaks and valleys in the backscattering amplitudes allow chemical identification of unknown elements and differentiation of different elements with sufficiently different $Z$ values, as well as interpolation or extrapolation for elements not calculated in the present study. The equations describing the peaks $\left(P_{i}\right)$ and valleys ( $V_{i}$ ) for the first $(i=1)$ and the second ( $i=2$ ) amplitude envelopes are

$$
\begin{align*}
& P_{1}=0.204(Z+8)  \tag{7a}\\
& V_{1}=0.136 Z  \tag{7b}\\
& P_{2}=0.136(Z-21) \tag{7c}
\end{align*}
$$

$$
\begin{equation*}
V_{2}=0.136(Z-39) \tag{7d}
\end{equation*}
$$

It is interesting to note that the slope of $P_{1}$ is 1.5 times the slopes of $P_{i}$ and $V_{i}$ where $i \neq 1$. Furthermore, with the exception of $P_{1}$, each peak and valley seem to start at the beginning of some series: $V_{1}$ with $Z=0, P_{2}$ with $Z=21$ (first transition series), and $V_{2}$ with $Z=39$ (second transition series).

Scatterer Phase. Figures $4 \mathrm{a}-\mathrm{c}$ depict the backscattering phase shifts for three groups while Figures $5 a-d$ show the scatterer phase variation for series of elements in the periodic table. These plots exhibit a considerable amount of scatter, especially at low $k$ and for the heavier elements like the rare earths. These scatters are systematic and are due to the truncation of the atomic wave function at the muffin tin radius $r_{\mathrm{mL}}$. Such a truncation introduces a discontinuity in the exchange and correlation potential at the muffin tin radius which in turn introduces oscillations with period $2 k r_{m \prime}$ in the phase shifts. It should be perfectly legitimate to smooth out such scatters in the tabulated phase shift before comparison is made with experiments. In practice, the scatter is small enough not to make too much difference.

At high enough $k$ values, $\phi_{b}(k)$ decreases almost linearly with increasing $k$, whereas at low $k$ values it exhibits complicated patterns which are related to the amplitude function $F(k)$. That is, above a certain energy, the phase shift decreases with increasing electron energy, whereas below such energy the phase shift depends heavily on the "resonance" interactions between the photoelectron and the various electronic shells of the scattering atom. A careful inspection of the low $k$ region of $\phi_{b}(k)$ suggests that the plateau (slow varying regions) and the inflection points (fast varying regions) in the scatterer phase correspond to the peaks and valleys, respectively, in the scattering amplitude. These are illustrated in Figures $4 \mathrm{a}-\mathrm{c}$ for three groups of elements where the $P_{i}$ arrows designate the plateau and the $V_{i}$ arrows represent the inflection points in $\phi_{b}(k)$. The $P_{i}$ and $V_{i}$ values (in $k$ ) are very similar to the corresponding values plotted in Figure 8.

It is interesting to note that the scatterer phase shift varies systematically with the atomic number $Z$. Within each shell, the phase shift increases linearly with $Z$ as a result of the increasingly positive potential. In Figure 9 we plot the scatterer phase shift for $k=P_{1}$ (solid curve) and $k=15.12 \AA^{-1}$ (dashed curve) as a function of atomic number $Z$. The former correspond to the phase shift at the photoelectron energy above which the phase shift decreases monotonically with increasing electron energy. It is immediately apparent that both curves vary linearly as a function of $Z$ with breaks (changing slope) at $Z \simeq 21$ and 57 which correspond to the starts of the $d$ and the $f$ shells, respectively. Small but significant periodic deviations from the linear curves occur (especially for $k=15.12$ $\AA^{-1}$ curve) due to the different electronic shells involved. Nevertheless, the $k=P_{1}$ curve can be used for chemical identification if the central atom phase shift (vide infra) can reliably be removed from the total phase shift.

The slope of the nearly linear or quadratic curve at $k \gtrsim P_{1}$ is also plotted as a function of $Z$ in Figure 9. Here it varies linearly with $Z$ only within each shell. The drastically different slopes for different shells produce discontinuities at the junctions. We have pointed out this phenomenon in our previous work on the parametrization of these phase functions (with $Z$ $\Sigma 35$ only). ${ }^{17 \mathrm{~b}}$

Central Atom Phase Shift. Central atom (or absorber) phase shift is a much simpler but perhaps stronger function of $k$. It generally decreases with increasing photoelectron energy (and hence $k$ ) as shown in Figures 6a-c for groups and Figures 7a-d for series of elements in the periodic table.

Again the central atom phase shift $\phi_{a}(k)$ varies systematically with increasing $Z$. In particular the potential is more


Figure 2. Backscattering arnplitude functions for (a) group 7A elements: (b) group 4A elements; (c) transition metals Fe . Ru. and Os.


Figure 3. Backscattering amplitude functions for some representative elements in (a) first transition series and beyond; (b) second transition series and beyond: (c) third transition series and beyond; (d) lanthanides.
attractive for increasing $Z$ and we expect that the phase shift should increase. This is indeed the case as exemplified in Figure 10, where we plot the central atom phase shifts for $k=3.78$, 9.45 , and $15.12 \AA^{-1}$ as a function of $Z$. It is apparent that all three curves are reasonably linear in $Z$ with, again, breaks at $Z \simeq 21$ and 57 which correspond to the injections of $d$ and $f$ electrons, respectively, into the electronic structure. The decreases in slope at these "break points" reflect the fact that the phase shift increases at slower rates for heavy elements than for light atoms. Once again, there are systematic deviations from these linear relations due to the different shells involved. Furthermore, the phase shifts at lower $k$ regions increase much
faster than the higher $k$ regions as $Z$ increases. At $Z \gtrsim 57$, however, all three curves approach similar slopes.

A comparison of the absorber phase shifts listed in Table 111 (calculated with Clementi-Roetti wave functions and the $Z$ +1 approximation) with those listed in Table V11 (calculated with Herman-Skillman wave functions) for the same elements indicates that the former is more positive than the latter. This is apparently due to the more positive potential introduced by the $Z+1$ approximation. Though the discrepancy between the two sets of phase shifts can in general be compensated by changing $E_{0}$, we recommend the use of Table VII whenever possible for consistency.

The $E_{0}$ Problem. It should be emphasized at this point that the phase shifts are unique only if the energy thresholds, $E_{0}$, are specified. Changing $E_{0}$ by $\Delta E_{0}=E_{0}{ }^{\prime}-E_{0}$ will change the momentum $k$ to

$$
\begin{equation*}
k^{\prime}=\left(k^{2}-\frac{2 \Delta E_{0}}{7.62}\right)^{1 / 2} \tag{8}
\end{equation*}
$$

where $k$ is in $\AA^{-1}$ and $\Delta E_{0}$ in eV . The corresponding modification of the phase shift function will be

$$
\begin{align*}
& \phi^{\prime}\left(k^{\prime}\right)=\phi(k)-2\left(k^{\prime}-k\right) r \\
& \simeq \phi(k)+2 r\left(\Delta E_{0}\right) / 7.62 k \tag{9}
\end{align*}
$$

for $2\left(\Delta E_{0}\right) / 7.62 \ll k^{2}$. As expected, the difference $\Delta \phi(k)=$ $\phi^{\prime}\left(k^{\prime}\right)-\phi(k)$ decreases with increasing $k$ indicating that phase shifts are more sensitive to a change in $E_{0}$ at small $k$ than at large $k$.

It is clear then that, in order to fit experimental data based upon some empirical $E_{0}$ with our theoretical phase shifts, we must allow $E_{0}$ to vary in such a way that $\Delta E_{0}=E_{0}{ }^{\text {th }}-E_{0}{ }^{\text {exp }}$ with $E_{0}{ }^{\text {th }}$ and $E_{0}{ }^{\exp }$ denoting the "theoretical" and "experimental" energy thresholds, respectively.

Since the determination of interatomic distance $r$ depends on the precise knowledge of $\phi(k)$, the nonuniqueness of phase shifts naturally causes concern about the uniqueness of the distance determination. Fortunately, it can be shown that by adjusting $E_{0}$ it is not possible to produce an artificially good fit with a wrong distance $r$, simply because changing $E_{0}$ will affect $\phi(k)$ mainly at low $k$ values by $\sim 2 r\left(\Delta E_{0}\right) /(7.26 k)$ whereas changing $r$ will affect $\phi(k)$ mostly at high $k$ values by $2 k(\Delta r) .{ }^{8}$

Effect of Electronic Configuration. In this work we chose to use the atomic ground state electronic configuration for most elements. For example, the valence shell configurations for groups 4A and 7A are $n s^{2} n p^{2}$ and $n s^{2} n p^{5}$, respectively. For transition metals, we use the "majority" configurations of
 third transition metal series, respectively, solely for the purpose of producing a smooth trend.

The small but significant effects of valence shell electronic configuration on the amplitude as well as the scatterer and central atom phase functions are illustrated in Figures $11 a-c$ and $12 \mathrm{a}-\mathrm{c}$ for Pd and Cu , respectively. The configurations used for Pd are $4 d^{8} 5 s^{2}, 4 d^{9} 5 s^{1}$, and $4 d^{10} 5 s^{0}$, whereas those used for Cu are $3 \mathrm{~d}^{9} 4 \mathrm{~s}^{2}$ and $3 \mathrm{~d}^{10} 5 \mathrm{~s}^{1}$. It can be seen that the amplitude functions (Figures 11a and 12a) are little affected by changes in electronic configuration. The small variations at low $k$ values are not unexpected because in this region the photoelectron energy is comparable to the valence shell binding energies. On the other hand, both the scatterer and the absorber phase shifts exhibit interesting systematic variations with electronic configuration. First, the scatterer phase (Figures 11 b and 12 b ) increases with increasing population of the $s$ orbital (or equivalently depopulation of the d orbitals). The difference between various configurations, however, diminishes as $k$ increases. For example, it decreases from 0.35 rad at $k=4 \AA^{-1}$ to 0.08 rad at $k=15 \AA^{-1}$ for Pd . As pointed out in the foregoing section, this difference, which is roughly inversely proportional to $k$, can largely be compensated for by changing $E_{0}$.

The effect of valence shell electronic configuration on central atom phase function follows the same trend. Figure 11 c depicts the $\phi_{a}{ }^{0}(k), \phi_{a}{ }^{1}(k)$, and $\phi_{a}{ }^{2}(k)$ functions for three different electronic configurations ( $4 \mathrm{~d}^{10-n} 5 \mathrm{~s}^{n}$ where $n=0,1,2$ ) of Pd while Figure 12c shows the $\phi_{a}{ }^{1}(k)$ functions for two different configurations ( $3 \mathrm{~d}^{10-n} 4 \mathrm{~s}^{n+1}$ where $n=0,1$ ) of Cu . In all cases, the phase shifts increase with increasing population of $s$ orbital (or increasing depopulation of d orbitals). The difference, however, diminishes with increasing $k$ (note that the

SCATTERER PHASE


SCATTERER PHASE


SCATTERER PHASE


Figure 4. Backscattering phase functions for (a) group 7A elements: (b) group 4A elements; (c) transition metals Fe. Ru. and Os. The arrows $P_{i}$ and $V_{1}$ designate the "plateau" and the "inflection" points in $\phi_{b}(k)$ which correspond to the "peaks" and "valleys" in $F(k)$ (cf. Figure 2).


Figure 5. Backscattering phase functions for some representative elements in (a) first transition series and beyond; (b) second transition series and beyond; (c) third transition series and beyond; (d) lanthanides.
near parallel appearance of the phase functions is an optical illusion). For example, it decreases from $\sim 0.25$ and 0.55 rad at $k=4 \AA^{-1}$ to 0.10 and 0.25 rad at $k=15 \AA^{-1}$ for Pd and Cu , respectively. Again, such a difference can largely be compensated by $E_{0}$ variation as shown in Figure 12c for copper.

Charge Effect. Throughout this paper we use neutral atomic entities in our calculations except for a few alkali or alkali-earth metals which are treated as cations.
The effect of atomic charge on central atom phase shift has been explored for several elements. A typical example is shown in Figure 13. Here we plot $\phi_{a}{ }^{0}, \phi_{a}{ }^{1}$, and $\phi_{a}{ }^{2}$ as a function of $k$ for Ca and $\mathrm{Ca}^{2+}$. It is immediately obvious that the dication
not only has a more positive phase shift but also a larger slope It is also readily apparent that the effect of atomic charge on phase shifts is significantly larger than that of electronic configuration. For example, at $k=4,15 \AA^{-1}, \phi_{a}{ }^{1}$ of Ca and $\mathrm{Ca}^{2+}$ are $0.20,-5.68$ and $2.64,-4.85$, respectively. This is not unexpected since the atomic charge exerts a significant effect on the central atom potential which is experienced by both the outgoing and the incoming photoelectrons. The difference, again, can partially be compensated by $E_{0}$ variation.

Comparison of $\phi_{a}{ }^{\prime}(\boldsymbol{I}=\mathbf{0}, 1,2)$ Functions. The $\phi_{a}{ }^{\prime}$ functions listed in Tables VI, VII, and VIII, where $/=0,1,2$, are central atom (absorber) phase shifts for the excitations $\mathrm{p} \rightarrow \mathrm{s}\left(L_{11,111}\right.$
edges), $\mathrm{s} \rightarrow \mathrm{p}$ ( $K$ or $L_{1}$ edge), and $\mathrm{p} \rightarrow \mathrm{d}$ ( $L_{11,111}$ edges), respectively. Two examples are shown in Figure 11 c for palladium and in Figure 13 for calcium. In both cases, there are large differences between the $\phi_{a}{ }^{\prime}$ phase functions. The order $\phi_{a}{ }^{2}>\phi_{a}{ }^{1}>\phi_{a}{ }^{0}$ as well as the divergence at large $k$ values may not be physically meaningful since subtracting $2 \pi$ from $\phi_{a}{ }^{2}$ and adding $2 \pi$ to $\phi_{a}{ }^{0}$ yield an inverted order $\phi_{a}{ }^{2}<\phi_{a}{ }^{1}$ $<\phi_{a}{ }^{0}$ which converges at high $k$.

We have also investigated the difference in absorber phase shifts for exciting 1s ( $K$ edge) vs. 2 s ( $L_{1}$ edge) and found no significant discrepancy between the two.

Relativistic Effect. We have obtained the charge density for tungsten calculated using relativistic wave functions from John H. Wood (Los Alamos). These were used as input to calculate the backscattering amplitude and phase. These are compared with the result using Herman-Skillman wave functions in Figure 14. We see that the amplitudes are in substantial agreement, especially beyond $k=6 \AA^{-1}$. The phase shows a systematic deviation which is progressively smaller for higher $k$. This kind of deviation is of the same order as that due to configuration differences and can be compensated for by changing $E_{0}$. Strictly speaking, for the heavier elements one should treat the electron scattering problem relativistically. ${ }^{27}$ However, the relativistic corrections such as spin-orbit terms are small compared with the Hartree and exchange potential and we assume can again be compensated for by $E_{0}$ change.

EXAFS Data Analysis. In EXAFS analysis, it is often necessary to multiply the $\chi(k)$ data by a weighting scheme such as $k^{n}$ in order to compensate for amplitude reduction as a function of $k$. For data in the range of $k=4-15 \AA^{-1}$, we recommend using the weighting schemes of $k^{3} \chi(k), k^{2} \chi(k)$, and $k \chi(k)$ for backscatterers with atomic number $Z \leqq 36,36 \leqq$ $Z \lesssim 57$, and $57 \lesssim Z \lesssim 86$, respectively. These weighting schemes appropriately emphasize the relative importance of different regions within the data without severely distorting the amplitude envelope. That is, for light scatterers with a rapidly attenuating backscattering amplitude (due partly to $F(k)$ and partly to the Debye-Waller factor) we suggest a larger $n$ value in $k^{n} \chi(k)$ to bring out the high $k$ region, whereas for heavy scatterers with strong backscattering power a small $n$ value will help preserve the "fine structure" of the backscattering amplitude as well as the backscattering phase. Another reason for choosing an appropriate weighting scheme is that the scattering amplitude exhibits increasing number of peaks (in $k$ space) with increasing atomic number $Z$ which inevitably show up in the Fourier transforms of the EXAFS spectra and could be mistaken as a separate distance or the satellite peak from a neighboring edge. In fact, the resulting Fourier transform and the filtered EXAFS spectra are to some extent affected by the weighting scheme and Fourier-filtering window. Experience showed that, with too narrow a filtering window and an inappropriate weighting, some of the "fine structures" of the amplitude envelope could be "washed" away.

Our theoretical functions and their parametrized versions ${ }^{17}$ have been used in EXAFS analysis of a wide variety of physical, chemical, and biological systems. The results are generally favorable. However, an overall scale factor of ca. $50 \%$ is often required in comparing the theoretical amplitude with experiment. ${ }^{28}$ For single-shell systems, the accuracy for the distance $r$ is better than $0.5 \%(\sim 0.01 \AA)$, for the Debye-Waller factor $\sigma$ is better than $10 \%$, and for the coordination number $N$ is $\sim 20 \% .^{15-20}$ Somewhat worse accuracy is expected for multiatom multidistance systems, though for systems with significantly different $Z$ scatterers interatomic distances can easily be determined to within $0.03 \AA .{ }^{16-20}$

Our previous work has been restricted to atoms with $Z<$ 36. In order to illustrate the situation for heavier elements, we compare the calculated $\mathrm{Pt}-\mathrm{Pt}$ phase shift and amplitude with

CENTRAL ATOM PHASE


CENTRAL ATOM PHASE


CENTRAL ATOM PHASE


Figure 6. Central atom (absorber) phase functions $\phi_{a}^{\prime}$ for (a) group 7A elements: (b) group 4A elements: (c) transition metals Ni, Pd, and Pt.


Figure 7. Central atom (absorber) phase functions $\phi_{a}{ }^{\prime}$ for some representative elements in (a) first transition series and beyond (Clementi-Roetti wave functions); (b) first transition series and beyond (Herman-Skillman wave functions); (c) second transition series and beyond; (d) third transition series, lanthanides, etc.
the first shell of platinum metal. Figure 15 a shows the $k \chi(k)$ data of platinum metal ( $L_{1}$ edge) after the "approximate" normalization with edge jump and Victoreen's true absorption coefficient $\mu_{0} / P=C \lambda^{3}-D \lambda^{4}(\text { where } C=470, D=219)^{29}$ and a cubic spline background removal with four sections. The energy threshold $E_{0}$ was arbitrarily chosen as 13900 eV , which corresponds to the first sharp peak at the edge. The Fourier transform is shown in Figure 15b. A smooth filtering window of 1.6-3.8 $\AA$ which encompasses the two major peaks ( 2.0 and $2.5 \AA$ ) was used to remove the higher shells as well as the high-frequency noise and the low-frequency residual peaks.

The Fourier peaks at 1.2 and $1.65 \AA$ (the former disappears when the data set is truncated at $k=3-14 \AA^{-1}$ ) may be due to residual EXAFS from the $L_{11}$ edge which is 608.9 eV lower in energy. The resulting Fourier filtered spectrum is shown as a dashed curve in Figure 15 a . Figure 15 c shows the comparison between the experimental amplitude and our theoretical function $N F(k) e^{-2 \sigma^{2} k^{2}}$ with an overall scale ${ }^{28}$ factor $N=0.54$ and a Debye-Waller factor $\sigma=0.05 \AA$ (note that $e^{-2 r / \lambda}$ was set to unity). The double peaks (ca. $0.5 \AA$ apart) observed for each shell are a consequence of the characteristic structure of the amplitude functions of heavy scatterers. The agreement


Figure 8. The positions (in $k$ space) of the peaks ( $P_{i}$ ) and valleys ( $V_{i}$ ) of the amplitude functions vs. atomic number $Z$.


Figure 9. Backscattering phase shift $\phi_{b}$ (rad) at $k=P_{1}$ (open circles) and $k=15.12 \AA^{-1}$ (open squares) vs. atomic number $Z$. The slope of $\phi_{b}$ for $k \leqq P_{1}$ is also plotted (filled circles) as a function of $Z$.


Figure 10. Central atom (absorber) phase shift $\phi_{a}{ }^{1}$ (radian) at $k=3.78$ (open circles), 9.45 (filled circles), and 15.12 (open squares) $\AA^{-1}$ vs. atomic number $Z$.


CENTRAL ATOM PHASE


Figure 11. Comparisons of the amplitude (a), backscattering phase (b), and central atom phase (c) functions for Pd with electronic configurations $4 d^{8} 5 s^{2}, 4 d^{9} 5 s^{1}$, and $4 \mathrm{~d}^{10} 5 s^{0}$.


Figure 12. Contrasts of the amplitude (a), backscattering phase (b), and central atom phase $\phi_{a}{ }^{1}$ (c) functions for Cu with electronic configurations $3 d^{9} 4 s^{2}$ and $3 \mathrm{~d}^{10} 4 \mathrm{~s}^{1}$. In (c), the dashed line corresponds to $\mathrm{d}^{10} \mathrm{~s}^{1}$ phase with $\Delta E_{0}=1.9 \mathrm{eV}$.

CENTRAL ATOM PHASE


Figure 13. The effect of atomic charge on the central atom phase shifts as exemplified by Ca and $\mathrm{Ca}^{2+}$.


Figure 14. Relativistic effects on the amplitude (a) and scatterer phase (b) functions as exemplified by a calculation on W using relativistic (dashed curves) vs. nonrelativístic (solid curves) wave functions.


Figure 15. (a) Unfiltered (solid curve) and Fourier filtered (dashed curve) $k \chi(k)$ vs. $k$ EXAFS data ( $L_{1}$ edge) of Pt metal after an "approximate" normalization with the edge jump and Victoreen's true absorption coefficient (see text) and a cubic spline background removal with four sections; (b) Fourier transform (solid curve) and the filtering window (dashed curve) of the unfiltered $k \chi$ ( $k$ ) vs. $k$ data; (c) experimental Pt backscattering amplitude function (solid curve) is fitted with the theoretical function (dashed curve) $N F(k) e^{-2 \sigma^{2} k^{2}}$ where $N=0.54$ and $\sigma=0.05 \AA$; (d) experimental $\mathrm{Pt}-\mathrm{Pt}$ phase shift (solid curve) is compared with the theoretical phase function (dashed curve) $\phi_{a b}=\phi_{a}+\phi_{b}-9 \pi$ where both $a$ and $b$ refer to Pt . In both (c) and (d), the theoretical energy threshold has been moved by 16 eV (viz., $E_{0}{ }^{\exp }-E_{0}{ }^{\mathrm{Th}}=16 \mathrm{eV}$ ).
is fairly good with the exception of low $k$ regions. While theoretical amplitudes are generally found to be too high at low $k$ values, it should be noted that the Fourier filtering with a reasonable window often causes a reduction of amplitude at low and high $k$ regions (cf. Figure 15a). Figure 15 d illustrates the good agreement between the experimental $\mathrm{Pt}-\mathrm{Pt}$ phase shift (after subtracting the $2 k R$ contribution where $R=2.775$ $\AA)^{30}$ and the theoretical function calculated by $\phi_{a b}=\phi_{a}+\phi_{b}$ $-9 \pi$ (where both $a$ and $b$ refer to platinum). It should be noted that the somewhat complicated structure (deviation from the almost quadratic $k$ dependence) of the phase function arises from the heavy scatterer platinum (cf. Figure 5c). It should also be mentioned that in Figures 15 c and d, the theoretical energy threshold has been moved by $16 \mathrm{eV}\left(\Delta E_{0}=\right.$ $\left.E_{0}{ }^{\text {exp }}-E_{0}{ }^{\prime \prime}\right)$. Good agreements have also been observed with $K$ and $L$ edges of other elements. ${ }^{31}$

In conclusion, our theoretical amplitude and phase functions are adequate for EXAFS data analysis for heavy elements with atomic number $Z \leqq 82$. These functions, or the parametrized versions, can be used not only to determine interatomic distances, Debye-Waller factors, and coordination numbers, but also to identify unknown chemical types in complicated systems, without resorting to searching, measuring, and analyzing model compounds. When coupled with the existing techniques such as Fourier transform and curve fitting, they greatly enhance the chemical content of EXAFS spectroscopy.

Acknowledgments. We thank A. L. Simons and B. Chambers for assistance in the programming.

## References and Notes

(1) R. de L. Kronig, Z. Phys., 70, 317 (1931); 75, 191, 468 (1932).
(2) E. A. Stern, Phys. Rev. B, 10, 3027 (1974).
(3) D. E. Sayers, F. W. Lyt(e, and E. A. Stern, Adv. X-Ray Anal. 13, 248 (1970).
(4) (a) D. E. Sayers, E. A. Stern, and F. W. Lytle, Phys. Rev. Lett, 27, 1204 (1971); (b) F. W. Lytle, D. E. Sayers, and E. A. Stern, Phys. Rev. B, 11, 4825 (1975); (c) E. A. Stern, D. E. Sayers, and F. W. Lytle, ibid., 11, 4836 (1975), and references cited therein.
(5) (a) D. E. Sayers, E. A. Stern, and F. W. Lytle, Phys. Rev. Lett., 35, 584 (1975); (b) D. E. Sayers, F. W. Lytle. M. Weissbluth, and P. Pianetta, J. Chem. Phys., 62, 2514 (1975); (c) F. W. Lytle, D. E. Sayers, and E. B. Moore, Jr., Appl. Phys. Lett., 24, 45 (1974).
(6) C. A. Ashley and S. Doniach, Phys. Rev. B, 11, 1279 (1975).
(7) P. A. Lee and J. B. Pendry. Phys. Rev. B, 11, 2795 (1975).
(8) P. A. Lee and G. Beni, Phys. Rev. B, 15, 2862 (1977)
(9) B. M. Kincaid and P. Eisenberger, Phys. Rev. Lett., 34, 1361 (1975).
(10) P. H. Citrin, P. Eisenberger, and B. M. Kincaid, Phys. Rev. Lett., 36, 1346 (1976).
(11) (a) B. M. Kincaid, P. Eisenberger, K. O. Hodgson, and S. Doniach, Proc. Natl. Acad. Sci. U.S.A., 72, 2340 (1975); (b) P. Eisenberger and B. M. Kincaid. Chem. Phys. Lett., 36, 134 (1975).
(12) P. Eisenberger and B. M. Kincaid, Sclence, 200, 1441 (1978).
(13) R. E. Watson and M. L. Perlman, Science, 199, 1295 (1978).
(14) (a) R. G. Shulman, P. Eisenberger, W. E. Blumberg, and N. A. Stombaugh, Proc. Natl. Acad. Sci. U.S.A., 72, 4003 (1975); (b) P. Eisenberger, R. G. Shulman, G. S. Brown, and S. Ogawa, ibid., 73, 491 (1976).
(15) R. G. Shulman, P. Eisenberger, B. K. Teo, B. M. Kincaid, and G. S. Brown, J. Mol. Biol., 124, 305 (1978).
(16) B. K. Teo, R. G. Shulman, G. S. Brown, and A. E. Meixner, submitted for publication.
(17) (a) B. K. Teo, P. A. Lee, A. L. Simons, P. Eisenberger, and B. M. Kincaid, J. Am. Chem. Soc., 99, 3854 (1977); (b) P. A. Lee, B. K. Teo, and A. L. Simons, ibid., 99, 3856 (1977).
(18) B. K. Teo, P. Eisenberger, and B. M. Kincaid, J. Am. Chem. Soc., 100, 1735 (1978).
(19) (a) B. K. Teo, K. Kijima, and R. Bau, J. Am. Chem. Soc., 100, 621 (1978); (b) B. K. Teo, P. Eisenberger, J. Reed, J. K. Barton, and S. J. Lippard, ibid., 100, 3225 (1978).
(20) (a) J. Reed, P. Eisenberger, B. K. Teo, and B. M. Kincaid, J. Am. Chem. Soc. 99, 5217 (1977); (b) J. Reed, P. Eisenberger, B. K. Teo, and B. M. Kincaid, ibld., 100, 2375 (1978).
(21) (a) S. P. Cramer, T. K. Eccles, F. Kutzler, K. O. Hodgson, and S. Doniach, J. Am. Chem. Soc.. 98, 8059 (1976); (b) S. P. Cramer and K. O. Hodgson, ibid., 100, 2748 (1978).
(22) For example, if one arbitrarily defines $\phi_{a}$ of atom $A$ (absorber), one can deduce $\phi_{b}$ of atom B (scatterer) from the experimental phase shift $\phi_{a b}$ for the atomic pair $A-B$. From $\phi_{b}$ one can then determine the central atom phase $\phi_{a}$ of any atom $A^{\prime}$ by measuring $\phi_{a^{\prime} b}$ which is the total phase shift of atom pair $A^{\prime} B$ where $A^{\prime}$ and $B$ denote the (new) absorber and the (old) scatterer, respectively. Similarly, from $\phi_{a}$ it is possible to deduce the scatterer phase $\phi_{b^{\prime}}$ of any atom $B$ by measuring $\phi_{a b^{\prime}}$ for the atom pair $A B$ with $A$ and $B$ being the (old) absorber and the (new) scatterer, respectively. All individual phase functions constructed in this manner are "relative" to the arbitrarily defined $\phi_{a}$ of absorber $A$
(23) S. M. Heald and E. A. Stern, Phys. Rev. B, 16, 5549 (1977).
(24) In the matrix element the initial state should be that of a neutral atom and the final state that of an ion with a $2 p$ hole. In our calculation, it is more
convenient to use either the neutral atom or the ion wave functions for both the initial and final states. The ratios $M_{21} / M_{01}$ obtained using these two methods are found to be in agreement to within a few percent even though the individual matrix elements show bigger variation. The result using the ion wave functions has been plotted in Figure 1. We should also mention that there is a considerable amount of literature dealing with $M_{21} / M_{01}$ for light elements or outer shells. See, for instance, O. J. Kennedy and S. T. Manson, Phys. Rev. A, 5, 227 (1972); K. Codling, R. G. Houlgate, J. B. West, and P. R. Woodruff, J. Phys, B, 9, L83 (1976).
(25) F. W. Lytle. D. E. Sayers, and E. A. Stern, Phys. Rev, B, 15, 2426 (1977).
(26) (a) E. Clementi and C. Roetti, At. Data Nucl. Data Tables, 14, 177 (1974); (b) F. Herman and S. Skillman, "Atomic Structure Calculations", Pren-tice-Hall, Englewood Cliffs, N.J., 1963.
(27) N. F. Mott, Proc. R. Soc. London. Ser. A, 124, 425 (1925); 135, 429 (1932): M. Fink and A. C. Yates, At. Data, 1, 385 (1970).
(28) The theoretical amplitude functions, which include inelastic processes in the scattering atom, are found to be off by $\sim 50 \%$. Part of this discrepancy is due to core relaxation effects. Furthermore, the amplitude is expected to be somewhat sensitive to the chemical environment and will depend on the distance $r_{i}$ (due to an exponential damping factor $e^{-2 r_{i} / \lambda}$ ). An overall scale factor is therefore included in the refinements
(29) 'International Tables for X-ray Crystallography', Vol. 1ll, Kynoch Press, Birmingham, England, 1968, pp 161, 172.
(30) L. Pauling. 'The Nature of the Chemical Bond', 3rd ed., Cornell University Press. Ithaca, N.Y., 1960, p 403.
(31) P. Rabe, G. Tolkiehn, and A. Werner, J. Phys. C, in press.

# Resonance Raman Spectroelectrochemistry. 6. Ultraviolet Laser Excitation of the Tetracyanoquinodimethane Dianion 

Richard P. Van Duyne, ${ }^{* 1 a, b}$ Mary R. Suchanski, ${ }^{1 c}$ Joseph M. Lakovits, ${ }^{1 a}$ Allen R. Siedle, ${ }^{1 \mathbf{d}}$ Keith D. Parks, ${ }^{1 \mathrm{a}}$ and Therese M. Cotton ${ }^{1 \mathrm{a}, \mathrm{e}}$<br>Contribution from the Department of Chemistry, Northwestern University, Evanston, Illinois 60201, and the Central Research Laboratories, 3M Company, St. Paul, Minnesota 55101. Received November 27,1978


#### Abstract

The resonance Raman spectrum has been obtained for the electrogenerated dianion of tetracyanoquinodimethane (TCNQ) upon excitation of its lowest energy electronic transition ( $\lambda_{\max } 330 \mathrm{~nm}$ ) with a frequency doubled, flashlamppumped, Rhodamine 640 dye laser. For comparison we report the normal Raman spectrum of solid $\mathrm{Li}_{2}$ TCNQ.THF. The electron transfer induced frequency shifts for the second reduction step of TCNQ are measured and interpreted using the $\pi$-bond order changes determined from SCF-MO-Cl and INDO/S electronic structure calculations as well as the $\pi$-bond length changes determined from a MNDO-SCF-MO calculation. Finally, the TCNQ ${ }^{2-}$ Raman data is used to identify the oxidation state of TCNQ in the coordination complex $\left[\mathrm{Co}(\text { acacen })(\mathrm{py})_{2}\right]_{2} \mathrm{TCNQ}$.


## Introduction

It is now widely recognized that the observables in resonance Raman spectroscopy (RRS) (viz., vibrational frequency, resonance-enhanced vibrational symmetry type, number and intensity pattern of overtones, and depolarization ratios) and their laser excitation wavelength dependence represent sensitive probes of the molecular and electronic structure changes that can occur in molecules. Such structure changes are commonly induced by chemical modification, electron-transfer (ET) reactions, and optical excitation. Our primary motivation for applying RRS to the study of molecular and electronic structure changes stems from a long-term interest in developing a detailed description of ET processes. In particular we have been concerned with evaluating the role of intramolecular vibrational energy dissipation processes in highly exothermic, homogeneous, ET reactions. ${ }^{2-7}$ To compare such ET theories with experiment, information is needed concerning the magnitude of the specific structural changes (viz., bond length, vibrational frequency, and anharmonicity) which occur within the donor and acceptor molecules during an ET process. In addition we are interested in studying the molecular and/or electronic structure changes that accompany the partial ET
reactions involved in the formation of donor-acceptor, charge-transfer complexes that behave as one-dimensional, organic, electrical conductors. ${ }^{8-12,41}$ Thus the technique of resonance Raman spectroelectrochemistry (RRSE) was developed ${ }^{13}$ as a convenient means of coupling the observational sensitivity of RRS for monitoring molecular and electronic structure changes with the ability of electrochemistry to initiate and cleanly carry out successive one-electron transfer reactions.

Tetracyanoquinodimethane (TCNQ) was chosen for study by RRSE because it is a strong electron-acceptor molecule, ${ }^{8}$ is the acceptor half of the prototype one-dimensional, organic metal tetrathiafulvalene-tetracyanoquinodimethane ${ }^{8,9}$ (TTF-TCNQ) and exhibits two successive, one-electron reductions that are both chemically and electrochemically reversible in deoxygenated, aprotic solvents: ${ }^{14}$

$$
\begin{gather*}
\operatorname{TCNQ}^{0}\left({ }^{1} \mathrm{~A}_{\mathrm{g}}\right)+\mathrm{e}^{-} \rightleftharpoons \operatorname{TCNQ}^{-} \cdot\left({ }^{2} \mathrm{~B}_{3 \mathrm{~g}}\right)  \tag{1}\\
\mathrm{TCNQ}^{-} \cdot\left({ }^{2} \mathrm{~B}_{3 \mathrm{~g}}\right)+\mathrm{e}^{-} \rightleftharpoons \operatorname{TCNQ}^{2-}\left({ }^{1} \mathrm{~A}_{\mathrm{g}}\right) \tag{2}
\end{gather*}
$$

RRSE with visible ion laser lines has been used to obtain the RRS of the ${ }^{2} \mathrm{~B}_{3 \mathrm{~g}}$ (viz., $D_{2 h}$ point group) ground state of the

